

## A Data Study of the Influence of the Equatorial Upper Stratosphere on Northern Hemisphere Stratospheric Sudden Warmings.

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### SUMMARY

Equatorial winds in the stratosphere are known to influence the frequency of stratospheric mid-winter sudden warmings. Sudden warmings, in turn, influence the Earth's climate both through their direct influence on polar temperatures and through the temperature dependence of ozone depletion in the lower stratosphere. The conventional (Holton-Tan) explanation for the equatorial influence on sudden warmings is in terms of the equatorial winds in the lower stratosphere ( $\sim 20$ - $30$  km) acting as a wave-guide for mid-latitude planetary wave propagation. This study employs stratospheric temperature analyses and equatorial rocketsonde wind data extending to 58 km to diagnose the relationship between the Northern Hemisphere polar temperatures and equatorial zonal winds at all height levels in the stratosphere. In addition to the recognised Holton-Tan relationship linking the polar temperatures to the quasi biennial oscillation in equatorial winds in the lower stratosphere, a strong correlation of polar temperatures with equatorial winds in the upper stratosphere is found. We suggest that this may be associated with the strength and vertical extent of the westerly phase of the semi annual oscillation in the upper stratosphere, although the observations alone cannot provide a conclusive, causal relationship. The main diagnostic tools employed are correlation studies and composite analysis. The results underline the need for continued high quality, equatorial wind measurements at all stratospheric levels.

KEYWORDS: Stratosphere Quasi biennial oscillation Semi annual oscillation Northern hemisphere winter Stratospheric sudden warming Inter-annual variability

### 1. INTRODUCTION

The Northern Hemisphere (NH) stratospheric winter circulation displays substantial inter-annual variability (Labitzke 1982). Some winters are extremely disturbed and are accompanied by midwinter stratospheric sudden warmings, in which the polar temperature increases by  $20^{\circ}\text{C}$  or more in just a few days (Andrews *et al.* 1987). In contrast, other winters have fewer warming events and the polar vortex remains cold and undisturbed. Variability associated with stratospheric warmings is the primary source of inter-annual variability in the NH winter lower stratosphere. The variability of NH winter stratospheric temperatures also appears to be closely coupled to the Arctic Oscillation, the leading mode of variability in the troposphere (Thompson and Wallace 1998, Baldwin and Dunkerton 1999).

Global annual-mean temperatures in the lower stratosphere have fallen by  $\sim 0.6^{\circ}\text{C}$  per decade over recent decades (e.g. WMO 1998). It is important to characterise natural variability if we are to assess how much of the observed temperature trend is directly related to human activity. It is also important because temperature crucially influences the chemical processes that lead to ozone trends in this region. Stratospheric warmings result in a weak, disturbed vortex with warm temperatures and substantial downward transport of ozone-rich air. Conversely, a strong, stable vortex with low temperatures will result in large chemical ozone loss (due to the temperature thresholds for polar stratospheric cloud formation and hence temperature dependence of chlorine activation) and weak transport of ozone (Chipperfield and Jones 1999). A number of studies have found evidence for long-term changes in the stratospheric circulation, including a strengthening

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and cooling of the Arctic vortex (Labitzke and van Loon 1995, Zurek *et al.* 1996). Any such changes would, for the reasons just stated, influence the long-term trends in ozone and temperature (Hood *et al.* 1997).

Although stratospheric sudden warmings have been observed and documented for many years, the factors that control their inter-annual variability are not well understood. An essential requirement for their development is the presence of quasi-stationary planetary wave disturbances. These waves are generated in the troposphere and propagate upward into the stratosphere, where the wave-1 component can be of sufficient amplitude to displace the vortex well away from the pole. Associated with each vortex displacement there is a rapid increase in temperature at the pole (whether or not the vortex is also disrupted).

The correlation between the winter-time average polar temperature and the phase of the quasi biennial oscillation (QBO) in the lower stratosphere was noted over twenty years ago (Holton and Tan 1980, 1982). The QBO is an oscillation of the equatorial zonal winds in the lower stratosphere, between easterlies and westerlies with an average period of 28 months. The maximum amplitude of the oscillation occurs around 25-30 km. When the equatorial zonal winds at 20-30 km are in an easterly QBO phase, the northern polar vortex is generally warmer, more disturbed by planetary waves and displacement and/or disruption of the vortex by midwinter warmings is more likely than in the westerly QBO phase (see also Dunkerton and Baldwin 1991, Baldwin and Dunkerton 1998). The conventional explanation for this correlation is that the QBO winds in the lower stratosphere influence the background mean flow, which then affects the propagation of planetary scale waves (e.g. McIntyre 1982; Holton and Austin 1991; O'Sullivan and Young 1992; O'Sullivan and Dunkerton 1994; Hamilton 1998; Niwano and Takahashi 1998; Baldwin *et al.* 2000). More specifically, the QBO winds in the lower stratosphere determine the position of the zero wind line near the equator, whose associated nonlinear critical layer acts as a wave-guide for the planetary wave propagation. This results in enhanced poleward heat transfer during an easterly QBO phase and weaker transfer during a westerly phase.

While this mechanism is generally accepted, it has also been recognised that the 'Holton-Tan' two-way distinction between 'warm disturbed easterly phase' and 'cold undisturbed westerly phase' does not always hold up, particularly in late winter (Hamilton 1998; Baldwin *et al.* 2000). Labitzke and van Loon (1988) and Naito and Hirota (1997) have noted that the periods when the Holton-Tan relation holds up well coincide with periods when the 11-year solar cycle is at its minimum phase (defined, for example, by the 10.7cm solar flux). Conversely, it appears to substantially weaken or reverse during periods of solar maximum. While the solar cycle / QBO link and the solar cycle / NH polar temperature links have been studied extensively (e.g. Labitzke and van Loon 1988; Gray and Dunkerton 1990; Teitelbaum and Bauer 1990; Salby and Shea 1991; Dunkerton and Baldwin 1992; Kodera 1991, 1993, 1995; Balachandran and Rind 1995; Rind and Balachandran 1995; Labitzke and van Loon 1996; Naito and Hirota 1997; Haigh 1999; Shindell *et al.* 1999), the precise mechanism that inter-relates (1) the QBO, which has its maximum amplitude in the lower-middle equatorial stratosphere ( $\sim$ 20-30 km), (2) the observed winter-time lower stratospheric polar temperatures and (3) the solar cycle, whose primary influence is in the region of the equatorial stratopause ( $\sim$ 40-50 km), has yet to be clearly elucidated.

In this paper, we explore the possible links between these three phenomena by examining the relationship between observed polar temperatures and two different equatorial zonal wind data sets. The first is the long time-series of wind observations from radiosonde ascents (Naujokat 1986) for the period 1955-99 that extend to approximately 32 km. The

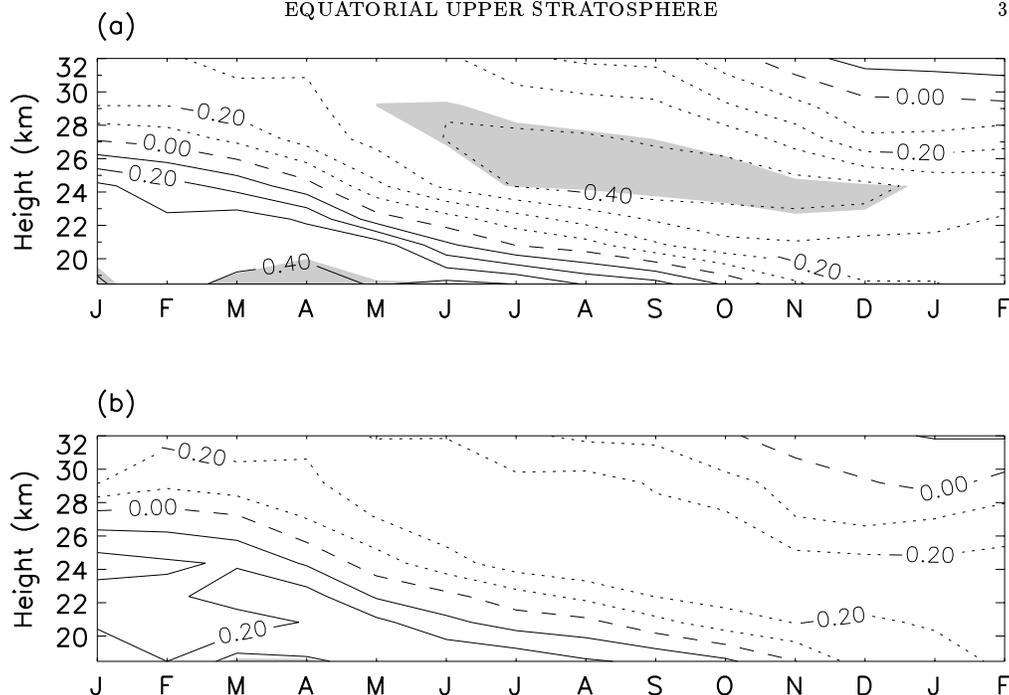


Figure 1. Correlation for the periods (a) 1964-90 and (b) 1955-99 between JF NP temperatures at 24 km and monthly-averaged  $u_{eq}$  from radiosonde ascents in each month from January of the previous year to February of the same year. Contour interval is 0.1. Dotted contours denote negative values. Dashed line is zero contour. Shading indicates values are significant at the 95% confidence level (no values were significant at the 99% confidence level). For details of the significance test applied, see appendix A.

second is a rocketsonde zonal wind data set for the period 1964-90 (Dunkerton and Delisi 1997; Dunkerton *et al.* 1998) that extends to 58 km. The rocketsonde dataset therefore includes not only the lower stratospheric winds (20-30 km) usually associated with the observed QBO modulation of the polar temperatures but also the higher altitudes that encompass the stratopause region, where the solar influence is greatest. In section 2 we briefly review the results of correlations between north polar temperatures and the lower stratospheric winds from the radiosonde dataset. In section 3 the rocketsonde dataset is described and a detailed analysis is presented of the relationships between north polar temperatures and equatorial winds at all heights to 58 km using the rocketsonde dataset. This analysis includes correlation studies, regression analyses, and composite studies. The main conclusions of the study are presented in section 4.

## 2. RADIOSONDE ANALYSIS

Figure 1a shows the correlation over 26 winters 1964 – 90 between the observed January-February average (JF) North Polar (NP) temperatures at 24 km and the equatorial zonal winds ( $u_{eq}$ ) from radiosonde observations in each month from January of the previous year to February of the same year. The NP temperatures are from the Berlin stratospheric analyses (courtesy K. Labitzke). The time interval 1964 – 90 is employed here to facilitate comparison with the rocketsonde analysis in the next section. We choose to use JF temperatures since most midwinter warmings occur in these months (and also to maintain consistency with previous studies). The NP temperatures are anti-correlated with the lower stratospheric  $u_{eq}$  in the previous months. The region of maximum correlation descends gradually through the atmosphere with time, corresponding to the descent

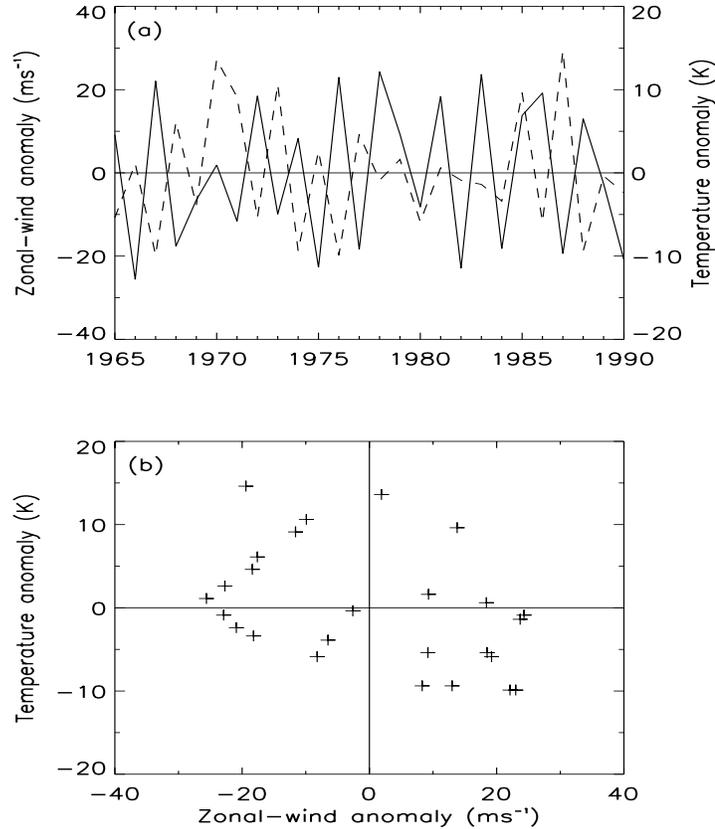


Figure 2. (a) Time-series of JF NP temperature anomalies (K) at 24 km ( $\sim 30\text{hPa}$ ) (dotted line) and  $u_{\text{eq}}$  anomalies ( $\text{ms}^{-1}$ ) at 24 km in the preceding December (solid line). (b) Scatter plot of the data.

of the QBO wind regimes. The sign of the correlation is consistent with the Holton-Tan mechanism, with easterly (negative)  $u_{\text{eq}}$  associated with warm (positive) polar temperatures and vice versa. For the 26-year period 1964–90, the maximum correlation coefficients are of order  $-0.4$ . However, when the correlation is repeated using data from 44 winters (1955–99, see figure 1b), which is the full extent of the available radiosonde data, the correlation drops in magnitude to only  $\sim -0.25$ .

Figure 2a shows the JF NP temperature anomalies at 24 km for the time interval 1964–90 (derived by subtracting the mean value over the period from each data point) and the  $u_{\text{eq}}$  anomaly at 24 km over the same time interval. It shows that the Holton-Tan relationship of negative correlation between the two signals holds for certain periods but not others. For example, the two signals are anti-correlated during the periods 1965–68, 1971–78 and 1986–88, but in the intervening years either the opposite is true (e.g. 1969–70, 1980–81 and 1984–85) or there is no apparent relationship (e.g. 1982–83). It is these latter periods that degrade the correlation, resulting in the relatively low correlations seen in figure 1. A scatter plot of the data (figure 2b) further demonstrates the weakness of the resulting correlation.

The difference in correlation displayed by the analysis of the two time intervals in fig. 1 can be understood in terms of the influence of the solar cycle. Figure 3 shows the time-series of 10.7 cm radio flux for 1955–99. Notice that the 1955–99 interval encom-

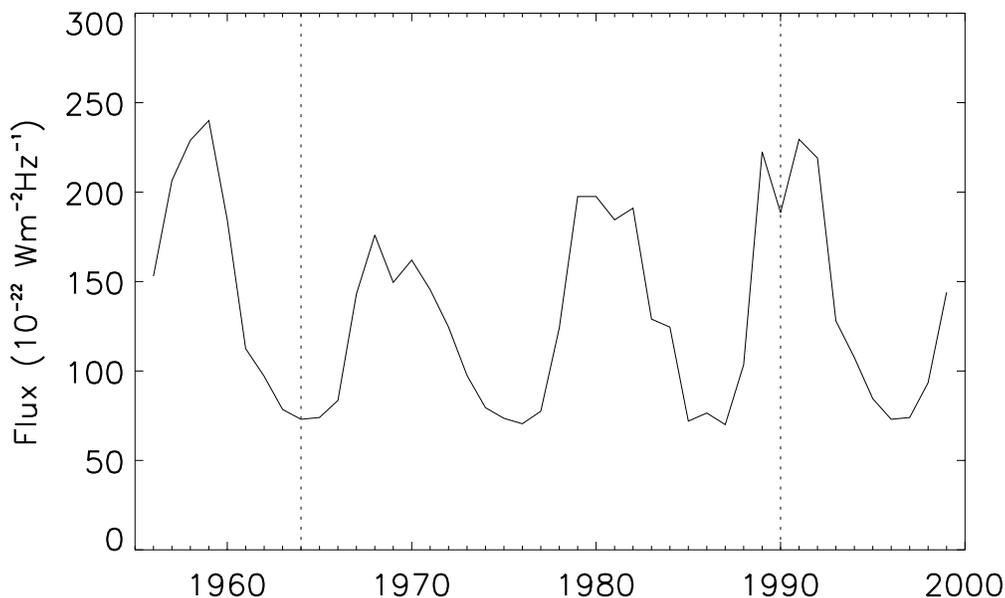


Figure 3. (a) Time-series of annual mean solar 10.7 cm radio flux ( $\times 10^{-22} \text{W m}^{-2} \text{Hz}^{-1}$ ). The dotted lines show the start and end of the period for which the rocketsonde data are available.

passes nearly four complete solar cycle periods while the 1964-90 interval encompasses only approximately 2.5 cycles with a slight bias towards solar minimum conditions (the solar maximum around 1968-70 is rather weak). Note also that the time interval 1962-78 originally examined in Holton and Tan's study was also biased towards solar minimum conditions (Naito and Hirota 1997). In figure 2a, the periods of Holton-Tan (i.e. negative) correlation tend to coincide with periods under solar minimum conditions and the periods in which the Holton-Tan relation is reversed (or absent) tend to coincide with periods under solar maximum conditions. It was Labitzke and van Loon (1988) who first noted this apparent interaction of the 11-year solar cycle with the QBO. In their study of the solar cycle influence on polar temperatures they separated the observed JF NP temperatures depending on the phase of  $u_{\text{eq}}$  in the lower stratosphere (see also Naito and Hirota 1997). In figure 4 we repeat their analysis for the interval 1955-99 to bring it up to date. The winters were classified using  $u_{\text{eq}}$  at 50 hPa (approximately 24 km) in accordance with the criterion used by Labitzke and van Loon. 'Transition' years when  $u_{\text{eq}}$  was within the range  $\pm 5 \text{ms}^{-1}$  were excluded. Note the strong positive correlation ( $r = +0.68$ ) of polar temperatures with the 10.7 cm flux in the west QBO phase and the negative correlation in the east QBO phase. The negative correlation appears to break down in the last few years of fig. 4b and this results in a correlation coefficient of only  $-0.14$  when the whole interval 1955-99 is included in the analysis. However, there was an extremely strong El-Nino in 1997 (Stockdale et al. 1998) which may have influenced the polar temperatures in that year.

In this study we take a slightly different approach to Labitzke and van Loon (1988) since we wish to explore the solar influence on the QBO / polar temperature relationship rather than the QBO influence on the solar / polar temperature relationship. We have therefore re-calculated the QBO / polar temperature correlations for the time interval

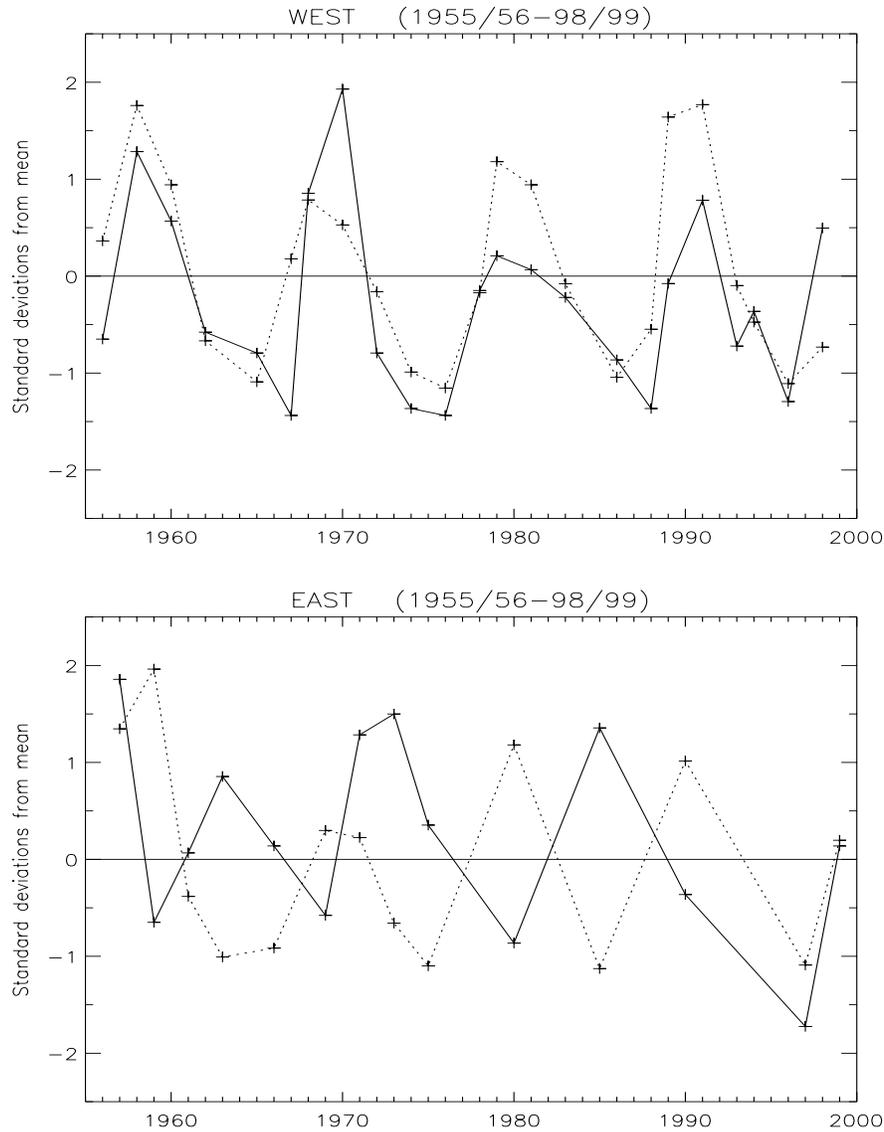


Figure 4. (a) Time series of NP temperature anomaly at 24 km ( $\sim 30$ hPa) (solid line) and 10.7 cm radio flux (dotted line). (a) west QBO phase years, (b) east phase years. Both signals have been plotted as standard deviations from the mean over the whole time-interval to aid comparison.

1955-99 shown in figure 1(b) after first separating the winters depending on whether the JF 10.7 cm solar flux was less than or greater than the mean for the whole period. Using this criteria we obtain 18 and 25 winters under solar maximum and solar minimum conditions, respectively. Under solar maximum conditions (figure 5a) there is a small positive correlation in the lower stratosphere suggesting that polar temperatures are warmer in the QBO westerly phase than in the QBO easterly phase i.e. in the opposite sense to the H-T relation. However, the amplitude of this positive correlation is extremely small. Under solar minimum conditions (figure 5b), on the other hand, there is a strong Holton-Tan relationship with a descending region of negative correlation reaching a maximum

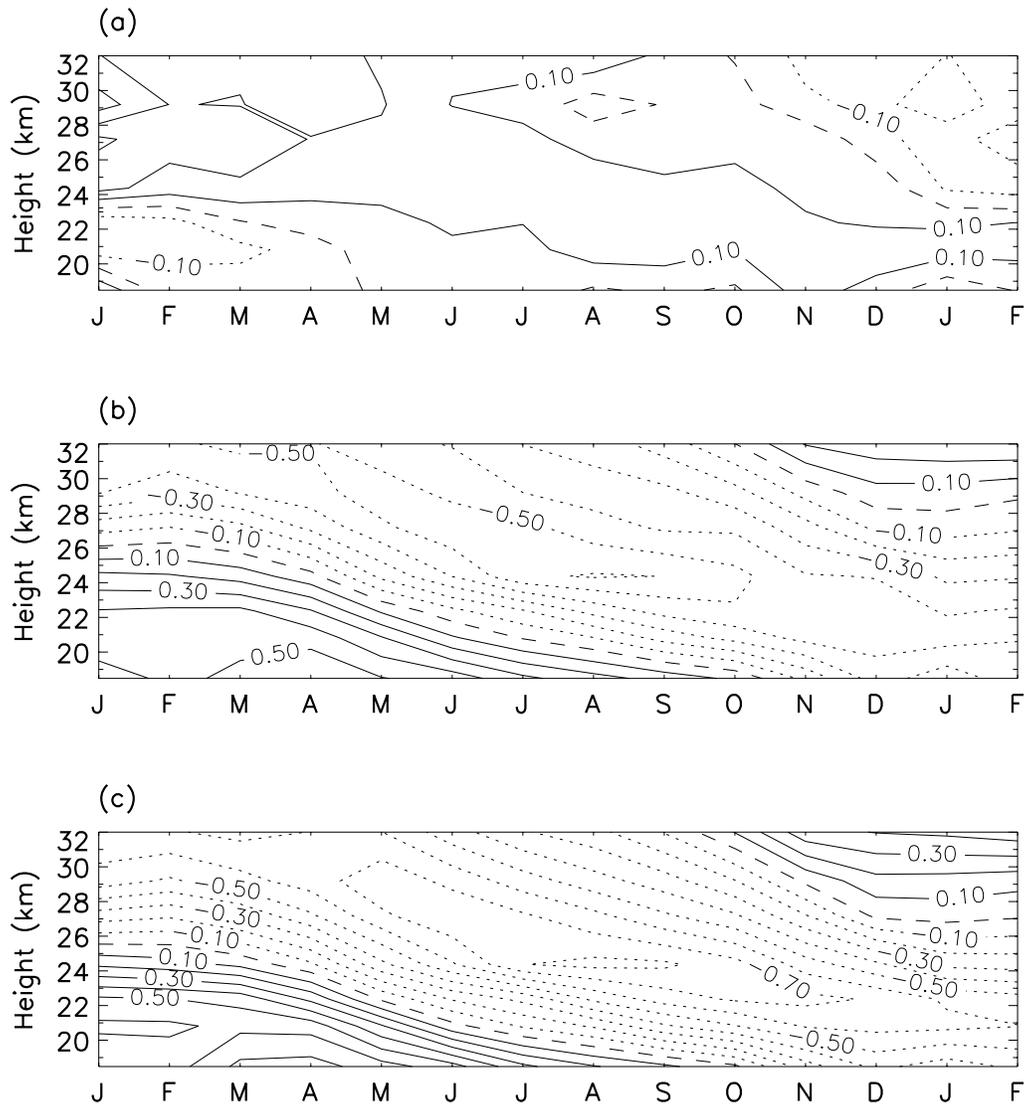


Figure 5. Correlation over the time interval 1955-99 between JF NP temperatures at 24 km and monthly-averaged  $u_{eq}$  from radiosonde ascents in each month from January of the previous year to February of the same year. (a) solar maximum years, (b) solar minimum years, (c) solar minimum years but for time interval 1955-96 only. Contour interval is 0.1. Dotted contours denote negative values. Dashed line is zero contour. Note that no confidence levels are estimated as the sampling interval varies in these stratified datasets and hence the autocorrelation of the data cannot be determined (see appendix A)

strength of  $-0.61$  at 24 km in August. If the analysis is repeated for the time interval 1955-96 in order to exclude the strong El-Nino event in 1997 this correlation increases in magnitude to  $-0.83$  (figure 5c).

### 3. ROCKETSONDE ANALYSIS

#### (a) *The data*

Regular rocketsonde ascents were made from Kwajalein ( $8^{\circ}\text{N}$ ) during the period 1969-90 and from Ascension Island ( $8^{\circ}\text{S}$ ) during the periods 1962-77 and 1979-89. Occasional gaps of up to a few months duration are present in both records. In order to obtain a single, unbroken time series of equatorial winds, the monthly-averaged data from the two stations were first de-seasonalised. This was done by calculating and subtracting the climatological monthly mean at each height and at each station. Gaps of up to three months duration in the records for each station were then filled by linear interpolation. For those heights and months when data were available for both stations, the two data sets were combined by averaging them together. When data were only available from one station, those data were used. In this way, an unbroken time series of de-seasonalised ‘proxy equatorial’ ( $8^{\circ}\text{S}$ – $8^{\circ}\text{N}$  averaged) zonal winds was achieved from January 1964 to February 1990 for the height region 20-58 km.

Figure 6a shows the time-height cross-section of the raw monthly-averaged data from Kwajalein. The QBO is the dominant signal in the lower stratosphere (20-35 km). The maximum amplitude of the oscillation is at around 25 km, with the wind varying between  $-30\text{ ms}^{-1}$  and  $+20\text{ ms}^{-1}$ . Above 35 km the semi-annual oscillation (SAO), with a period of six months, dominates. Figure 6b shows the combined Kwajalein and Ascension Island data set, which was derived as described above. (Note that the SAO has been removed by the deseasonalising process). The QBO period has been further highlighted in figure 6c using a 9-60 month bandpass filter.

#### (b) *Correlation studies*

The correlation study to examine the link between  $u_{\text{eq}}$  and JF NP temperatures was repeated using the rocketsonde wind observations to extend the analysis into the upper equatorial stratosphere. Figure 7 shows the correlation at each height and month leading up to January/February (i.e. identical to figure 1a but using the rocketsonde data). Note first the negative correlation of  $\sim 0.4$  at 20 – 35 km, which slowly propagates downwards with time. This was the correlation noted by Holton and Tan nearly 20 years ago and below 30 km it is virtually identical to the results of the correlations using radiosonde wind data (figure 1a).

However, there is also significant correlation at higher levels. In the 30 – 50 km region there is a positive correlation of  $\sim 0.3 - 0.4$ , also descending with time. This change in sign of the correlation with height reflects the next incoming phase of the equatorial wind QBO (see figure 6). Above 50 km, there is a region of relatively high negative correlation in September/October. At 52 km in September, the correlation reaches  $-0.6$  and is significant at the 99% confidence level. We note here that there can be no suggestion that this high correlation is simply a response of  $u_{\text{eq}}$  to the winter polar warming, since the strongest correlation is found between the JF polar temperatures and  $u_{\text{eq}}$  in the previous September, 4-5 months before the midwinter warmings take place.

The rocketsonde correlation analysis was also extended to investigate the possible solar cycle influence, although the limited timespan of the rocketsonde data (just over two solar cycle periods) should be borne in mind in the interpretation of these results. As before, the winters were separated depending on whether the JF 10.7 cm radio flux was less than or greater than the mean for the period, giving 12 and 14 winters in solar minimum and solar maximum conditions respectively. The results (figure 8) are consistent with the radiosonde analysis (figure 5) in showing stronger correlations under solar minimum conditions than under solar maximum conditions. The peak correlation

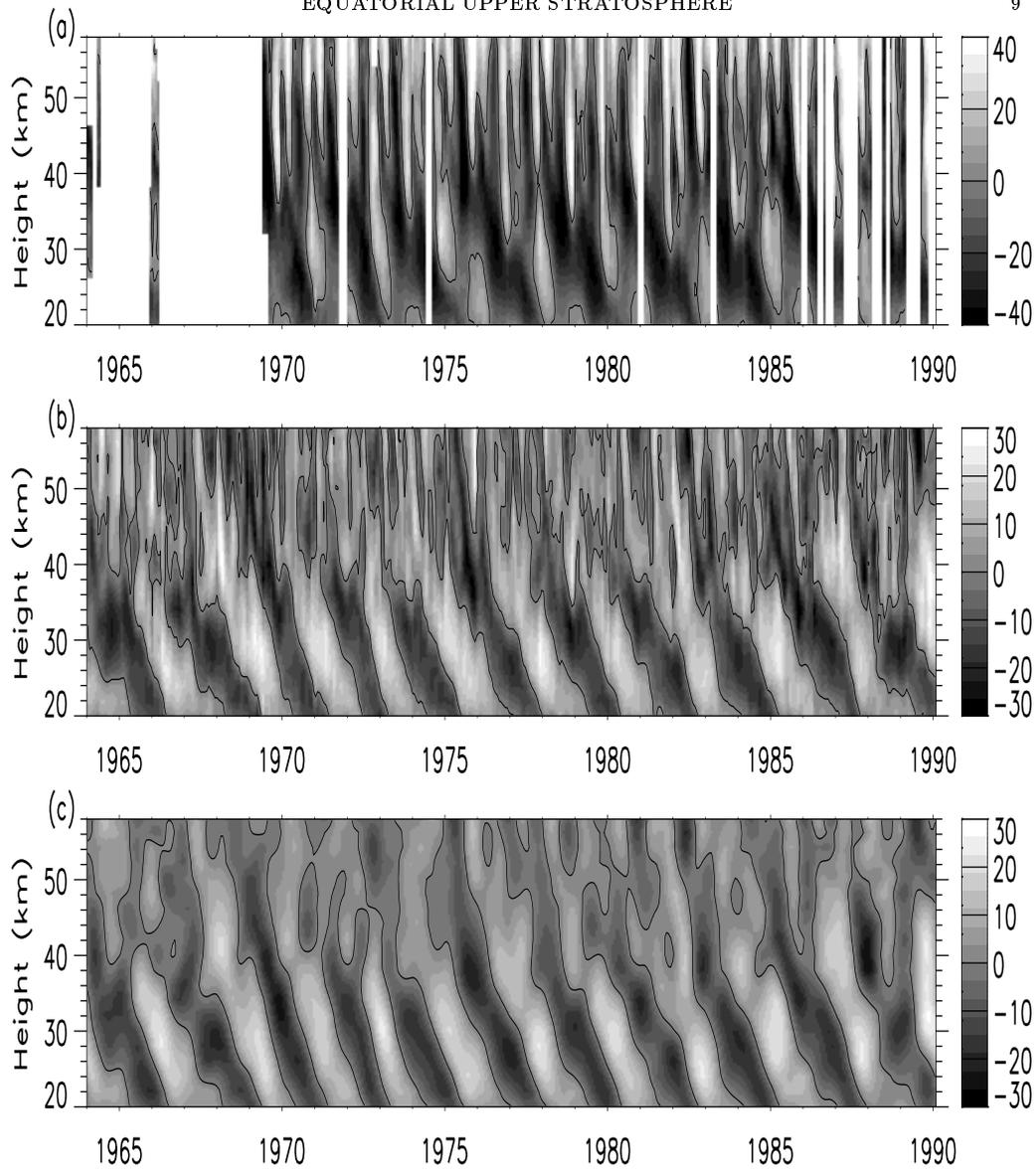


Figure 6. (a) Monthly-averaged zonal wind ( $\text{ms}^{-1}$ ) derived from rocketsonde ascents at Kwajalein ( $8^{\circ}\text{N}$ ). (b) De-seasonalised monthly-averaged zonal wind dataset derived by combining data from both Kwajalein and Ascension Island as described in the text. (c) As in panel (b) after passing through a 9-60 month bandpass filter. Zero value is denoted by contour line.

amplitudes under solar minimum conditions are  $-0.80$  at 24 km in October,  $+0.75$  at 44 km in August and  $-0.75$  at 52 km in September.

(c) *Regression analyses*

Linear regression analyses were carried out in order to assess the effectiveness of the rocketsonde  $u_{\text{eq}}$  at different heights and in different months as a predictor of the winter-time polar temperature. Initially, simple linear regression was used to fit the JF NP temperature to  $u_{\text{eq}}$  at (a) 24 km in December (as an indicator of the phase of the

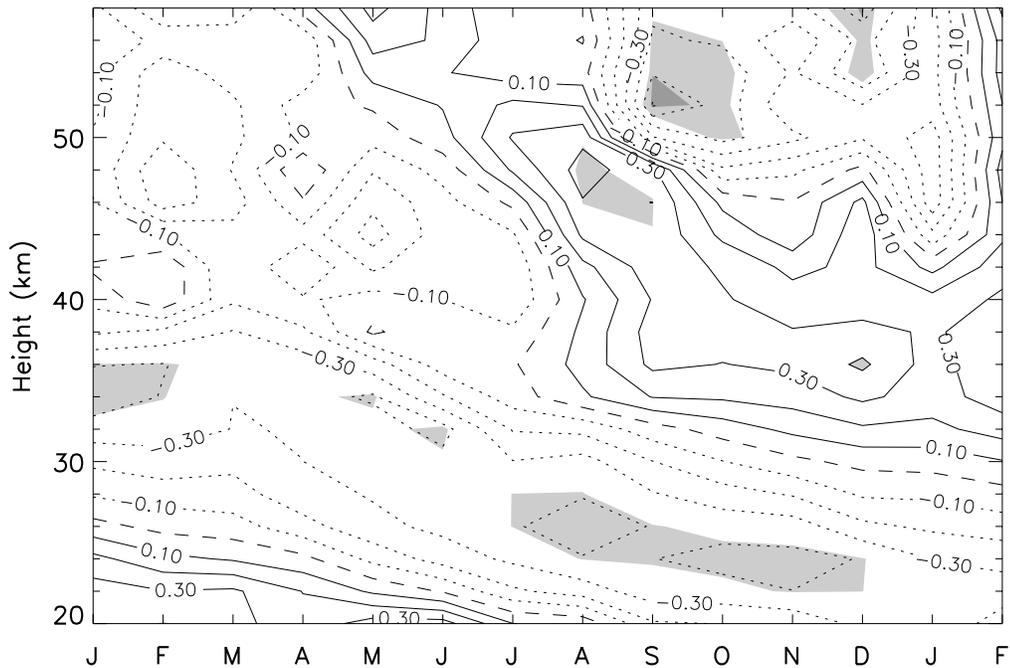


Figure 7. Correlation for the period 1964-90 between JF NP temperatures at 24 km and monthly-averaged  $u_{eq}$  from rocketsonde ascents in each month from January of the previous year to February of the same year. Contour interval is 0.1. Dotted contours denote negative values. Dashed line is zero contour. Shading indicates values are significant at the 95% confidence level (no values were significant at the 99% confidence level). For details of the significance test applied, see appendix A.

QBO in the lower stratosphere), and (b) 52 km in September. The best fit to  $u_{eq}$  at 52 km was found to describe 37% of the variance in the polar temperature. By contrast, the best fit to  $u_{eq}$  at 24 km was found to describe only 16% of the variance.

A series of multiple linear regressions was then carried out in order to find the two parameters that best predict the polar temperatures. In each analysis,  $u_{eq}$  for a given height and month was employed as a fixed parameter, while  $u_{eq}$  for every other available height and month was used in turn as the second parameter. In this way, we sought to maximise the amount of the variance in the polar temperature that could be described. In the first analysis,  $u_{eq}$  at 24 km in December was employed as the fixed parameter. It was found that  $u_{eq}$  at 52 km in September was most effective as the second parameter, with a total of 41% of the variance in the polar temperature being described.

Repeating the analysis with  $u_{eq}$  at 52 km in September as the fixed parameter, it was found that the most effective second parameter was  $u_{eq}$  at 48 km in the same month. A total of 51% of the variance in the polar temperature was described. This is an interesting result on two accounts. Firstly, the greatest additional contribution did not come from the lower stratosphere. Secondly, the two parameters consist of  $u_{eq}$  immediately above and immediately below 50 km in September. In figure 7 it can be seen that there is a sharp change in the sign of the correlation at this point. This suggests that the strength of the vertical gradient of  $u_{eq}$  at  $\sim 50$  km in September may be a useful predictor of the NP temperature in the following winter.

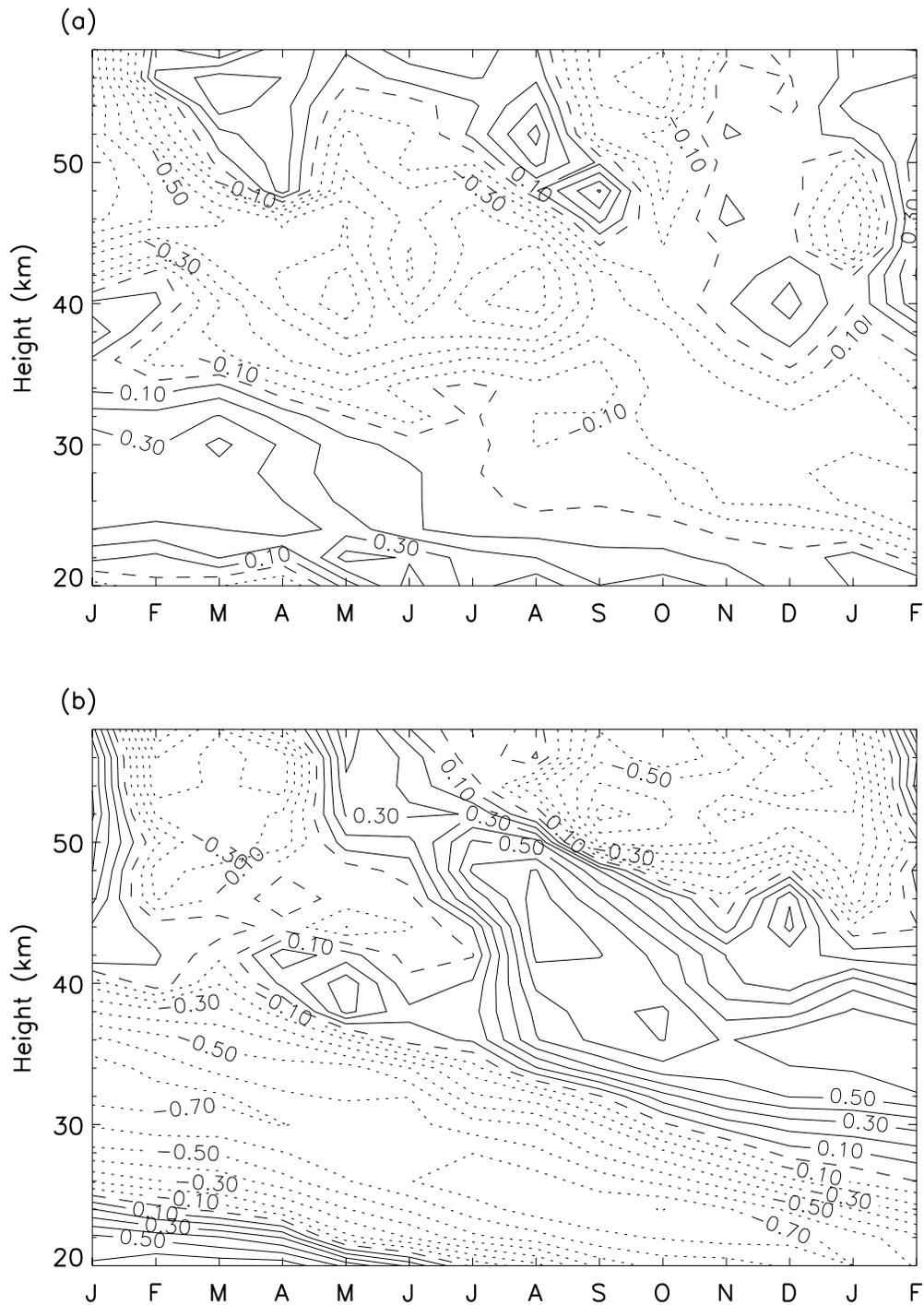


Figure 8. Correlation for the period 1964-90 between JF NP temperatures at 24 km and monthly-averaged  $u_{eq}$  from rocketsonde ascents in each month from January of the previous year to February of the same year. (a) solar maximum conditions, (b) solar minimum conditions. Contour interval is 0.1. Dotted contours denote negative values. Dashed line is zero contour. Note that no confidence levels are estimated as the sampling interval varies in these stratified datasets and hence the autocorrelation of the data cannot be determined (see appendix A)

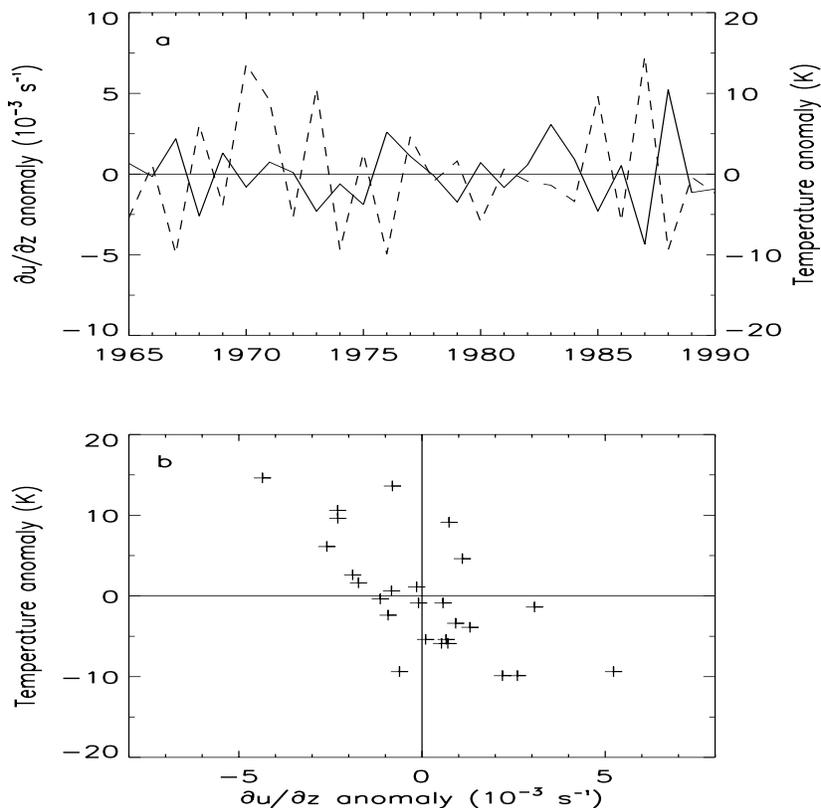


Figure 9. (a) Time-series of JF NP temperature anomalies (K) at 24 km ( $\sim 30$ hPa) (dotted line) and  $\partial u_{eq}/\partial z$  anomalies ( $\times 10^{-3} s^{-1}$ ) at 50 km in the preceding September (solid line). (b) Scatter plot of the data.

In order to investigate this possible link further,  $\partial u_{eq}/\partial z$  was derived from the raw rocketsonde data using centred differences. These values were then de-seasonalised and combined in the same manner as that described above for the equatorial winds. The correlation between the JF NP temperature and  $\partial u_{eq}/\partial z$  was found to peak at 50 km in September, with a correlation of -0.68. Using simple linear regression it was found that  $\partial u_{eq}/\partial z$  at 50 km in September could describe 46% of the variance in the JF NP temperature. The multiple linear regression described above was repeated, using  $\partial u_{eq}/\partial z$  at 50 km in September as the fixed parameter and all possible values of both  $u_{eq}$  and  $\partial u_{eq}/\partial z$  as the second parameter. It was found that  $u_{eq}$  at 44 km in October was most effective as the second parameter, with a total of 62% of the variance in the NP temperatures being described.

These results suggest a link between the JF NP temperatures and  $u_{eq}$  and its vertical gradient in the upper stratosphere in the preceding September/October. This strong link is evident in figure 9a, which shows the JF NP temperatures for 1965-90 overlaid by  $\partial u_{eq}/\partial z$  at 50 km in the preceding September. The two signals are reasonably well anti-correlated over most of the period. The years that are less well anti-correlated (e.g. 1971, 1977-8) appear to be randomly distributed and, unlike in figure 2a, do not display a solar cycle pattern. Unfortunately, the length of the rocketsonde data interval is too short to produce meaningful results if the above analysis is repeated after first separating

into solar minimum and solar maximum conditions. A comparison of the scatter plot of the data (figure 9b) with the corresponding plot in figure 2b for the lower stratospheric winds ( $u_{eq}$  at 24 km) clearly demonstrates that the relationship between the JF polar temperature and  $\partial u_{eq}/\partial z$  at 50 km is stronger than with  $u_{eq}$  at 24 km.

(d) *Composites*

In figure 10 we show a series of latitude - height ‘composite differences’ of JF winds using daily zonally-averaged zonal wind data derived from the National Centers for Environmental Prediction / National Center for Atmospheric Research (NCEP/NCAR) 1200 UTC heights and temperatures (Kalnay et al. 1996). Firstly, composites were derived by averaging together the JF wind distributions from the five warmest and five coldest Northern Hemisphere winters between 1964/65 and 1989/90. The appropriate years were selected on the basis of the NP (Berlin) temperatures at 24 km. We do this to highlight the differences in the JF wind distribution between years that exhibit a major warming and those that do not. The ‘cold’ composite minus the ‘warm’ composite is shown in figure 10a. During cold winters the westerly polar vortex is less disturbed by planetary waves and their accompanying warming events and hence there is a strong, high-latitude westerly anomaly, with maximum composite differences of  $35 \text{ ms}^{-1}$  at 32 km ( $\sim 10 \text{ hPa}$ ). There is a small westerly anomaly at the equator, consistent with the Holton-Tan mechanism, peaking at around  $11 \text{ ms}^{-1}$ .

Secondly, the wind distributions were composited on the basis of  $u_{eq}$  at 24 km in January (using the rocketsonde data, which are virtually identical to the radiosonde data at this level). These composites test the Holton-Tan mechanism whereby the polar temperatures and vortex are influenced by  $u_{eq}$  in the lower stratosphere. The ‘west QBO’ minus ‘east QBO’ composite differences in figure 10b show a high latitude westerly wind anomaly peaking at  $\sim 15 \text{ ms}^{-1}$  at 32 km, consistent with a stronger westerly polar jet during the westerly phase of the QBO at 24 km.

Finally, composites were constructed on the basis of  $\partial u_{eq}/\partial z$  at 50 km in September (from the rocketsonde data). The difference between the ‘large shear’ and ‘small shear’ years has a maximum amplitude of  $\sim 25 \text{ ms}^{-1}$  at high latitudes (figure 10c). Hence, years with stronger than average vertical wind shear near the equatorial stratopause in the autumn months tend to be followed by a cold, more stable westerly polar vortex, and vice versa. These composites confirm that, while by no means accounting for the full amplitude of the differences in figure 10a, the composite differences at high latitudes based on  $\partial u_{eq}/\partial z$  at 50 km in September are substantially larger than those constructed using  $u_{eq}$  at 24 km.

In figure 11 we show the height-time evolution of  $u_{eq}$  from the rocketsonde data for the ‘warm’ and ‘cold’ composites described above, and the difference between them. The major feature between 40-60 km is the semi-annual oscillation (SAO), in which  $u_{eq}$  reverses direction from easterly to westerly and back again on a time-scale of approximately 6 months. September is the onset of a westerly phase at around 50 km. There is a notable difference in the nature of this westerly phase between the ‘warm’ and ‘cold’ composites. In the ‘warm’ composite, the westerly phase extends deeper in the atmosphere, with the zero wind line at its lowest extent in November reaching  $\sim 30 \text{ km}$  compared with only  $\sim 40 \text{ km}$  in the ‘cold’ composite. This is a result of a more rapid descent of the westerly phase during September and October. The vertical wind shear at  $\sim 50 \text{ km}$  in September reflects this difference, being weaker in the ‘warm’ composite than in the ‘cold’ composite (as we have already noted in the discussion of figure 10c above).

Figure 11 suggests that the difference between the warm and cold composites may be associated with the relative strength and depth of the SAO westerly phase around

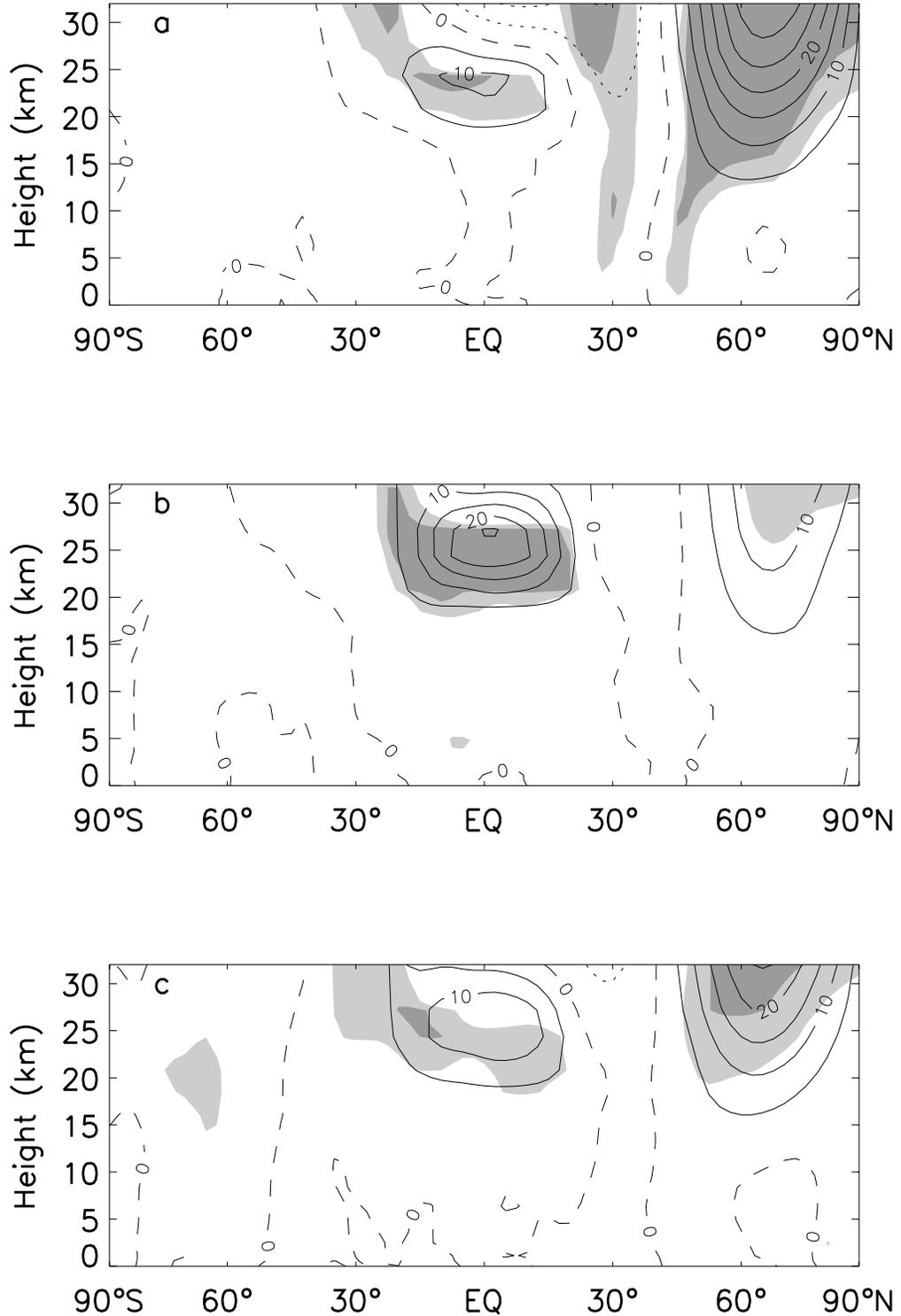


Figure 10. Latitude-height distributions of JF zonally-averaged zonal wind ( $\text{ms}^{-1}$ ) from the NCEP analyses for the period 1964-90. (a) composite difference of the 5 coldest minus 5 warmest NH winters as determined by the 30 hPa JF NP temperatures, (b) composite difference of 5 most westerly minus 5 most easterly QBO phase based on  $u_{\text{eq}}$  at 24 km in the previous December, (c) composite difference of 5 strongest minus 5 weakest  $\partial u_{\text{eq}}/\partial z$  at 50 km in the previous September. Contour interval is  $5 \text{ ms}^{-1}$ . Negative (easterly) values are denoted by dotted contours. Light and dark shading denote regions where values are significant at the 95% and 99% confidence levels respectively. The significances of the values were tested by performing a Student's *t*-test on the difference between the two mean wind distributions.

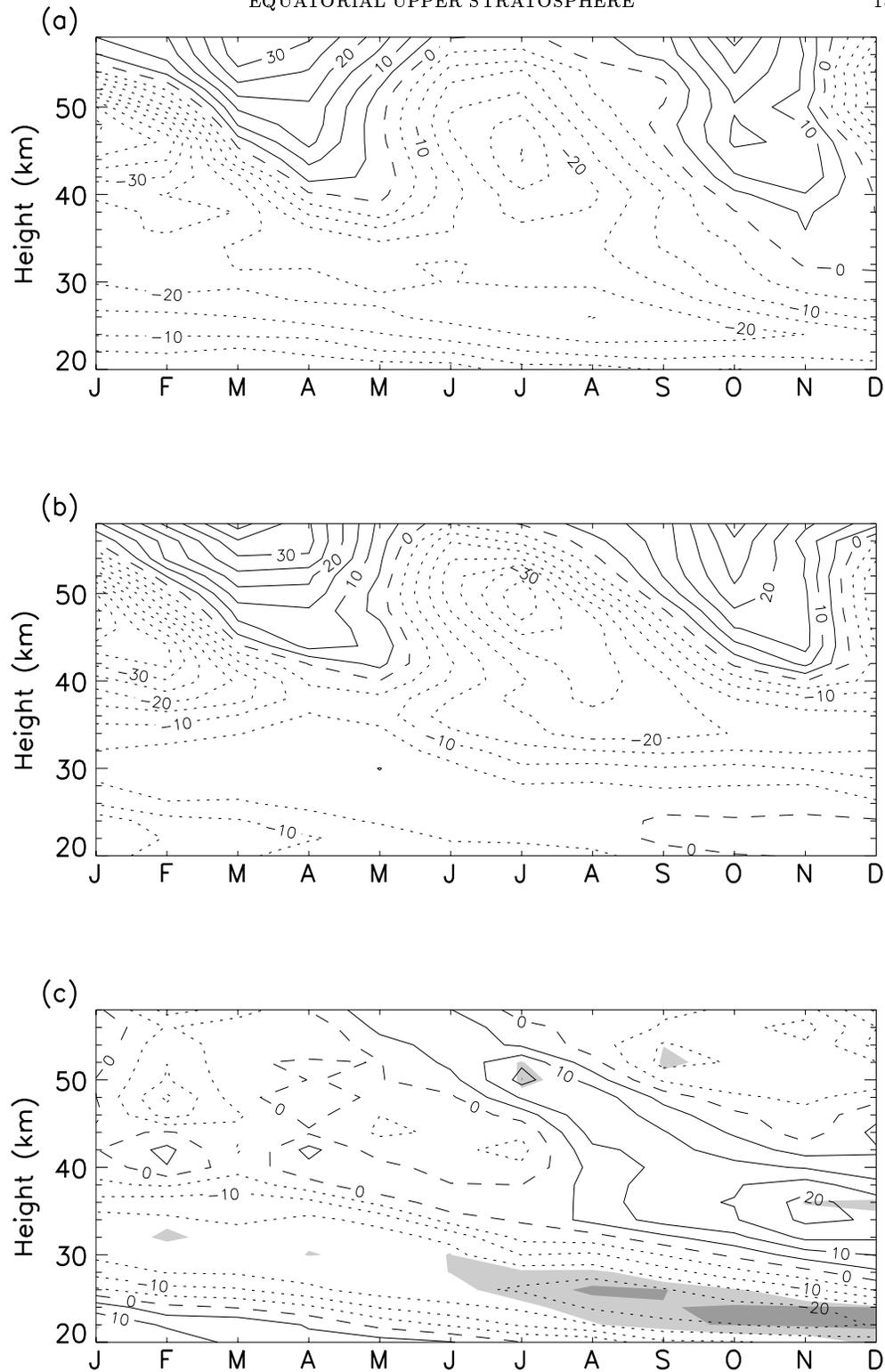


Figure 11. Time-series of monthly-averaged  $u_{eq}$  ( $\text{ms}^{-1}$ ) measured by rocketsonde. (a) Composite of 5 warmest NH winters, as determined using 30 hPa NP temperatures, (b) As for panel (a) but for the 5 coldest winters. (c) The difference between the two composites i.e. panel (a) minus panel (b). Contour interval is  $5 \text{ ms}^{-1}$ . Negative (easterly) values are denoted by dotted contours. Light and dark shading denote regions where values are significant at the 95% and 99% confidence levels respectively. The significances of the values were tested by performing a Student's t-test on the difference between the two mean wind distributions.

September/October. The westerly phase of the SAO arises as a result of vertically propagating westerly-phase equatorial waves e.g. gravity and Kelvin waves (Andrews *et al.* 1997). These waves are able to propagate relatively easily through lower stratospheric easterlies. However, when the waves reach heights ( $\sim 50$  km) where the background zonal winds are of similar magnitude and direction to their own phase speeds, the waves are unable to propagate further. They are dissipated and their Eliassen-Palm fluxes converge, accelerating the background flow, thus forcing it even more westerly. In this way, the westerly SAO phase gradually descends through the atmosphere with time. The rate of descent of the westerly phase, and hence the vertical wind shear at 50 km, is therefore dependent on the QBO in the lower stratosphere, since the relative ease of vertical propagation of the westerly-phase-speed waves depends on the wind speed and direction in the lower stratosphere (Lindzen and Holton, 1968; Gray and Pyle 1989; Hirota *et al.* 1991; Kennaugh *et al.* 1997). Strong easterlies at 20-30 km allow enhanced penetration of such waves and hence more effective westerly forcing and more rapid descent (and vice versa). The resulting QBO modulation of the zonal wind at these upper levels has been observed in the HRDI wind observations (Burrage *et al.* 1996) and is evident in figure 6c which shows a weak QBO signal extending into the upper stratosphere. It is for this reason, presumably, that we see evidence of a QBO-like variation in  $\partial u_{\text{eq}}/\partial z$  at 50 km in figure 9 and a fairly substantial westerly anomaly in the lower equatorial stratosphere in the composite difference of figure 10c, similar to that in figure 10b.

#### 4. CONCLUSIONS

We have presented evidence for a strong correlation between January / February north polar (NP) temperatures and the equatorial wind ( $u_{\text{eq}}$ ) and its vertical shear in the upper stratosphere in the region of the stratopause ( $\sim 50$  km). This suggests an influence on polar winter temperatures from the upper stratosphere, in addition to that from the lower stratospheric QBO, which is already well documented. This is perhaps not surprising given that the planetary waves believed to be primarily responsible for stratospheric warmings are deep structures that encompass the whole depth of the stratosphere. The equatorial upper stratosphere is also the part of the stratosphere where the solar cycle has a significant influence. Separating the winters into solar minimum and solar maximum conditions enhances the derived correlation under solar minimum conditions but weakens it under solar maximum conditions (figure 5). The maximum correlation between the NP temperatures and  $u_{\text{eq}}$  occurs in the previous September / October (figure 7), at the onset of a westerly phase of the SAO. Warm winters appear to be characterised by a westerly SAO phase that is weaker in amplitude and penetrates deeper into the lower stratosphere than in the colder winters (figure 11). However, September and October are fairly quiescent months and it may be inappropriate to place undue emphasis on these months. It may be that the mechanism linking the equatorial upper stratosphere and polar temperatures is present throughout the winter months but is less evident later on because of the more disturbed nature of the winter months.

While a data correlation study such as this cannot, on its own, establish a causal relationship nor provide the mechanism for this relationship (if, indeed, one exists) it nevertheless suggests that a better understanding of the equatorial upper stratosphere and its influence on to the polar stratosphere, will be essential if we are to fully understand the inter-annual variability of NH polar temperatures. The study also highlights the need for continued high quality wind measurements of the equatorial upper atmosphere. The 1990s has been an unusual decade, with very few midwinter warmings. This may be

an indicator of climate change. Unfortunately, our rocketsonde analysis could not be extended into the 1990s because of the lack of data.

#### APPENDIX

In figures 1 and 7 the significance of each point on the correlation plot was tested by calculating the value of the statistic  $t = r\sqrt{\frac{N_e-2}{1-r^2}}$ , where  $r$  is the correlation between the two time-series and  $N_e$  is the number of independent data points in the time-series. In the null case of no correlation, this statistic is distributed like Student's  $t$ -distribution with  $N_e - 2$  degrees of freedom (Press et al. 1992). In order to compensate for the autocorrelation of the wind data, the effective sample size  $N_e = N$  where  $N$  is the actual number of data points in the time-series unless  $\rho_1 > 0$  where  $\rho_1$  is the lag-one autocorrelation coefficient, in which case  $N_e = N \left\{ \frac{1-\rho_1}{1+\rho_1} \right\}$  (Wilks 1995).

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