

1 Effective degrees of nonlinearity in a family of generalized models of two-dimensional turbulence

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We study the small-scale behavior of generalized two-dimensional turbulence governed by a family of model equations, in which the active scalar $\theta = (-\Delta)^{\alpha/2} \psi$ is advected by the incompressible flow $\mathbf{u} = (-\psi_y, \psi_x)$. Here ψ is the stream function, Δ is the Laplace operator, and α is a positive number. The dynamics of this family are characterized by the material conservation of θ , whose variance $\langle \theta^2 \rangle$ is preferentially transferred to high wave numbers (direct transfer). As this transfer proceeds to ever-smaller scales, the gradient $\nabla \theta$ grows without bound. This growth is due to the stretching term $(\nabla \theta \cdot \nabla) \mathbf{u}$ whose “effective degree of nonlinearity” differs from one member of the family to another. This degree depends on the relation between the advecting flow \mathbf{u} and the active scalar θ (i.e., on α) and is wide ranging, from approximately linear to highly superlinear. Linear dynamics are realized when $\nabla \mathbf{u}$ is a quantity of no smaller scales than θ , so that it is insensitive to the direct transfer of the variance of θ , which is nearly passively advected. This case corresponds to $\alpha \geq 2$, for which the growth of $\nabla \theta$ is approximately exponential in time and nonaccelerated. For $\alpha < 2$, superlinear dynamics are realized as the direct transfer of $\langle \theta^2 \rangle$ entails a growth in $\nabla \mathbf{u}$, thereby, enhancing the production of $\nabla \theta$. This superlinearity reaches the familiar quadratic nonlinearity of three-dimensional turbulence at $\alpha = 1$ and surpasses that for $\alpha < 1$. The usual vorticity equation ($\alpha = 2$) is the border line, where $\nabla \mathbf{u}$ and θ are of the same scale, separating the linear and nonlinear regimes of the small-scale dynamics. We discuss these regimes in detail, with an emphasis on the locality of the direct transfer.

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I. INTRODUCTION

The production of progressively smaller scales, possibly to be limited by viscous effects only, in incompressible fluid flow at high Reynolds numbers, is a fundamental problem in fluid dynamics. This long-standing problem is of genuine interest for obvious reasons. One is that the production of small scales plays a key role in the possible development of singularities from smooth initial conditions in the three-dimensional (3D) Euler or Navier-Stokes equations that govern the flow. Another reason is that in the presence of a large-scale forcing, a persistent production of small scales would be crucial to maintain a spectral energy flux (direct energy cascade). The realizability of such a steady and viscosity-independent flux is central to the Kolmogorov theory of turbulence as this would be required to rid the virtually inviscid energy inertial range of the injected energy, thereby, making it possible for a statistical equilibrium to be established. This dynamical scenario is either explicitly or implicitly assumed to apply to other fluid systems as well, not just the 3D Navier-Stokes equations. For example, in the Kraichnan-Batchelor [1–3] theory of two-dimensional (2D) turbulence, the dynamics of the mean-square vorticity (twice the enstrophy) are assumed to be synonymous in many aspects to those of the 3D energy. In particular, the enstrophy injected into the system at large scales is hypothesized to cascade to a dissipation range at small scales. As another example, the mean-square potential vorticity in the quasigeostrophic geophysical flow model is believed to behave in a similar manner [4]. Thus “cascading dynamics” have been considered universal among fluid systems.

The evolution of fluid flow is intrinsically nonlinear because of the quadratic advection term, which couples all scales of motion. Apparently, this is an underpinning reason

for the cascade universality mentioned in the preceding paragraph. However, the “effective degree of nonlinearity” of the small-scale dynamics is not always quadratic and differs from one system to another. The implication is that the presumed cascades would have fundamental differences and would not be universal in a strict sense. For an example of the discrepancy in the effective degree of nonlinearity among fluid systems, let us consider the respective evolution equations for the 3D vorticity $\boldsymbol{\omega}$ and 2D vorticity gradient $\nabla \omega$ given by

$$\partial_t \boldsymbol{\omega} + (\mathbf{u} \cdot \nabla) \boldsymbol{\omega} = (\boldsymbol{\omega} \cdot \nabla) \mathbf{u}, \quad \nabla \cdot \mathbf{u} = 0 \quad (1)$$

and

$$\partial_t \nabla \omega + (\mathbf{u} \cdot \nabla) \nabla \omega = \boldsymbol{\omega} \mathbf{n} \times \nabla \omega - (\nabla \omega \cdot \nabla) \mathbf{u}, \quad \nabla \cdot \mathbf{u} = 0, \quad (2)$$

where \mathbf{u} is the fluid velocity and \mathbf{n} is the normal to the fluid domain in 2D. The stretching term $(\boldsymbol{\omega} \cdot \nabla) \mathbf{u}$ for the 3D vorticity $\boldsymbol{\omega}$ in Eq. (1) is essentially quadratic in $\boldsymbol{\omega}$ because the velocity gradient $\nabla \mathbf{u}$ is expected to behave as $\boldsymbol{\omega}$ on phenomenological grounds. As a consequence, an explosive 3D vorticity growth from a smooth initial vorticity field is possible, if not inevitable [5,6]. In contrast, the stretching term $(\nabla \omega \cdot \nabla) \mathbf{u}$ for the 2D vorticity gradient $\nabla \omega$ in Eq. (2) is virtually linear in $\nabla \omega$ because $\nabla \mathbf{u}$ is well behaved in the sense that the mean-square vorticity $\langle \omega^2 \rangle = \langle |\nabla \mathbf{u}|^2 \rangle$ is conserved. (Note that the rotation term $\boldsymbol{\omega} \mathbf{n} \times \nabla \omega$ does not affect the amplitude of $\nabla \omega$.) As a result, the growth of 2D vorticity gradients can possibly be approximately exponential in time only, a relatively mild behavior. Hence, one would expect profound differences between the (highly nonlinear) 3D vorticity and the (nearly linear) 2D vorticity gradient dynamics. A notable example of these differences is that in the inviscid

86 limit, the 2D enstrophy dissipation rate vanishes [7–9],
 87 whereas the 3D energy dissipation rate presumably remains
 88 nonzero. Another example is the discrepancy in the depen-
 89 dence on the Reynolds number of the number of degrees of
 90 freedom in the two cases [10].

91 The effective degree of nonlinearity in the above sense
 92 differs not only between 2D and 3D fluids but also among
 93 2D fluid systems. In this study, we investigate this varying
 94 degree among members of a broad family of generalized
 95 models of 2D turbulence, first introduced by Pierrehumbert
 96 *et al.* [11]. By doing so, we extend several previous studies
 97 [12–15], aiming to unify our understanding of turbulent
 98 transfer in physically realizable fluid systems. The family’s
 99 dynamics are characterized by the material conservation of
 100 the active scalar $\theta = (-\Delta)^{\alpha/2}\psi$, whose variance $\langle \theta^2 \rangle$ is prefer-
 101 entially transferred to high wave numbers (small scales).
 102 Here ψ is the stream function, Δ is Laplace’s operator, and α
 103 is a positive number. As the transfer of $\langle \theta^2 \rangle$ proceeds to
 104 ever-smaller scales, the gradient $\nabla\theta$ grows without bound.
 105 This growth is due to the stretching term $(\nabla\theta \cdot \nabla)\mathbf{u}$, whose
 106 effective degree of nonlinearity depends on α and is wide
 107 ranging, from approximately linear to highly superlinear.
 108 Linear behavior is realized when $\nabla\mathbf{u}$ is a quantity of no
 109 smaller scales than θ , so that the transfer of $\langle \theta^2 \rangle$ to the small
 110 scales (direct transfer) has no significant effects on $\nabla\mathbf{u}$. In
 111 other words, θ behaves nearly passively. This case corre-
 112 sponds to $\alpha \geq 2$, for which $\nabla\theta$ can grow approximately ex-
 113 ponentially in time without acceleration. For $\alpha < 2$, superlin-
 114 ear dynamics can be realized as the direct transfer of $\langle \theta^2 \rangle$
 115 entails a growth in $\nabla\mathbf{u}$, thereby, enhancing the production of
 116 $\nabla\theta$. This superlinearity reaches the familiar quadratic nonlin-
 117 earity of three-dimensional turbulence at $\alpha=1$ and exceeds
 118 that for $\alpha < 1$. The usual vorticity equation ($\alpha=2$) is the bor-
 119 der line, where $\nabla\mathbf{u}$ and θ are of the same scale ($\langle |\nabla\mathbf{u}|^2 \rangle$
 120 $= \langle \theta^2 \rangle$), separating the linear and nonlinear regimes of the
 121 small-scale dynamics. We discuss these dynamical regimes
 122 in detail, with an emphasis on the local nature of the transfer
 123 of $\langle \theta^2 \rangle$. The implication of the present results is that a com-
 124 prehensive theory for this family of generalized 2D turbu-
 125 lence needs to account for the wide range of effective de-
 126 grees of nonlinearity of the family’s small-scale dynamics.

127 II. GOVERNING EQUATIONS

128 The equation governing the evolution of the family of
 129 active scalars $\theta = (-\Delta)^{\alpha/2}\psi$ (for $\alpha > 0$) advected by the in-
 130 compressible flow $\mathbf{u} = (-\psi_y, \psi_x)$ is

$$131 \quad \theta_t + \mathbf{u} \cdot \nabla\theta = 0. \quad (3)$$

132 This equation was proposed by Pierrehumbert *et al.* [11] in
 133 an attempt to better understand the nature of transfer locality
 134 in 2D turbulence, by examining how turbulent transfer re-
 135 sponds to changes in the parameter α . Equation (3) is physi-
 136 cally relevant for selected values of α . The usual 2D vorticity
 137 equation corresponds to $\alpha=2$. When $\alpha=1$, Eq. (3) is known
 138 as the surface quasigeostrophic equation and governs the ad-
 139 vection of the potential temperature, which is proportional to
 140 $\theta = (-\Delta)^{1/2}\psi$, on the surface of a quasigeostrophic fluid. In
 141 addition to the genuine interest due to this physical signifi-

142 cance [12–20], the surface quasigeostrophic equation has re-
 143 ceived some special attention for its resemblance to the 3D
 144 Euler system [21–24]. A mathematical feature of particular
 145 interest is the possible development of finite-time singulari-
 146 ties (from smooth initial conditions), which, as argued by
 147 pioneering studies [21,22,25] of this problem, could be asso-
 148 ciated with the formation of weather fronts in the atmo-
 149 sphere. This, however, appears not to be the case [26].

150 For simplicity, we consider Eq. (3) in a doubly periodic
 151 domain of size L , and all fields concerned are assumed to
 152 have zero spatial average. This allows us to express the
 153 stream function as

$$\psi(\mathbf{x}, t) = \sum_{\mathbf{k}} \hat{\psi}(\mathbf{k}, t) \exp\{i\mathbf{k} \cdot \mathbf{x}\}. \quad (4)$$

154 Here $\mathbf{k} = 2\pi L^{-1}(k_x, k_y)$, where k_x and k_y are integers not si-
 155 multaneously zero. The reality of ψ requires $\hat{\psi}(\mathbf{k}, t) = \hat{\psi}^*($
 156 $-\mathbf{k}, t)$, where the asterisk denotes the complex conjugate. The
 157 fractional derivative $(-\Delta)^{\alpha/2}$ (which can be readily extended
 158 to $\alpha < 0$, though not considered in this study) is defined by
 159

$$\begin{aligned} \theta(\mathbf{x}, t) &= (-\Delta)^{\alpha/2}\psi(\mathbf{x}, t) = \sum_{\mathbf{k}} k^\alpha \hat{\psi}(\mathbf{k}, t) \exp\{i\mathbf{k} \cdot \mathbf{x}\} \\ &= \sum_{\mathbf{k}} \hat{\theta}(\mathbf{k}, t) \exp\{i\mathbf{k} \cdot \mathbf{x}\}, \end{aligned} \quad (5)$$

160 where $k = |\mathbf{k}|$ is the wave number. Equation (3) expresses ma-
 161 terial conservation of θ , which gives rise to an infinite set of
 162 conserved quantities. In particular, the generalized enstrophy
 163 (active scalar variance)
 164

$$Z = \frac{1}{2} \langle \theta^2 \rangle = \frac{1}{2} \langle |(-\Delta)^{\alpha/2}\psi|^2 \rangle = \frac{1}{2} \sum_{\mathbf{k}} k^{2\alpha} |\hat{\psi}(\mathbf{k}, t)|^2 \quad (6)$$

165 is conserved. In addition, the generalized energy
 166

$$E = \frac{1}{2} \langle \psi\theta \rangle = \frac{1}{2} \langle |(-\Delta)^{\alpha/4}\psi|^2 \rangle = \frac{1}{2} \sum_{\mathbf{k}} k^\alpha |\hat{\psi}(\mathbf{k}, t)|^2 \quad (7)$$

167 is also conserved. Note that E is the usual kinetic energy
 168 when $\alpha=2$, while Z is the usual kinetic energy when $\alpha=1$.
 169 Only for these cases is the kinetic energy conserved. The
 170 modal powers (spectra) of E and Z differ by the factor k^α .
 171 Therefore, the redistribution of a non-negligible amount of E
 172 to small scales would violate the conservation of Z . Simi-
 173 larly, the redistribution of a non-negligible amount of Z to
 174 large scales would violate the conservation of E . This means
 175 that if a spectrally localized profile is to spread out in wave-
 176 number space, most of E and Z get transferred to large and
 177 small scales, respectively. This is the basis for the dual cas-
 178 cade hypothesis in 2D turbulence. Here we are mainly con-
 179 cerned with the direct transfer of Z . A more complete treat-
 180 ment should include the inverse transfer of E as well since
 181 these are known to be intimately related.
 182

183 Given Eq. (4), we can express $\mathbf{u} = (-\psi_y, \psi_x)$ in terms of a
 184 Fourier series in the form
 185

$$u(\mathbf{x}, t) = i \sum_{\mathbf{k}} (-k_y, k_x) \hat{\psi}(\mathbf{k}, t) \exp\{i\mathbf{k} \cdot \mathbf{x}\}. \quad (8)$$

186

187 By substituting Eqs. (5) and (8) into Eq. (3), we obtain the
 188 evolution equation for each individual Fourier mode $\hat{\theta}(\mathbf{k}, t)$
 189 $= k^\alpha \hat{\psi}(\mathbf{k}, t)$ of the conserved quantity θ

$$\frac{d}{dt} \hat{\theta}(\mathbf{k}, t) = \sum_{\ell+m=\mathbf{k}} \frac{(m^\alpha - \ell^\alpha) \ell \times \mathbf{m}}{\ell^\alpha m^\alpha} \hat{\theta}(\ell, t) \hat{\theta}(\mathbf{m}, t), \quad (9)$$

190

191 where $\ell \times \mathbf{m} = \ell_x m_y - \ell_y m_x$. The sum on the right-hand side of
 192 Eq. (9) involves all modes [except $\hat{\theta}(\mathbf{k}, t)$] and is a measure
 193 of the level of “excitation” of the mode $\hat{\theta}(\mathbf{k}, t)$ due to all
 194 admissible wave vector triads $\mathbf{k} = \ell + \mathbf{m}$. For a given triad, the
 195 coupling coefficient $(m^\alpha - \ell^\alpha) \ell \times \mathbf{m} / (\ell^\alpha m^\alpha)$ depends on α .
 196 Its magnitude, together with the magnitudes of the coupling
 197 coefficients in the governing equations for $\hat{\theta}(\ell, t)$ and
 198 $\hat{\theta}(\mathbf{m}, t)$, is a measure of triad dynamical activity, in the sense
 199 that larger (in magnitude) coupling coefficients correspond to
 200 more intense modal dynamics. This is intimately related to
 201 the effective degree of nonlinearity and locality of the small-
 202 scale dynamics as will be seen in the subsequent sections.

203 III. EFFECTIVE DEGREES OF NONLINEARITY OF THE 204 SMALL-SCALE DYNAMICS

205 We now examine the behavior of $\nabla \theta$. Generally speaking,
 206 any derivative $(-\Delta)^\eta \theta$, for $\eta > 0$, can be called a small-scale
 207 quantity. Here we consider $\nabla \theta$, which is a “twin brother” of
 208 $(-\Delta)^{1/2} \theta$, for its special status in Eq. (3) as well as its math-
 209 ematical tractability. For $\alpha = 2$, a similar treatment of $\Delta \theta =$
 210 $-\Delta \omega$ can be carried out in the same manner.

211 A. Growth of the active scalar gradient

212 The governing equation for $\nabla \theta$ is

$$\partial_t \nabla \theta + (\mathbf{u} \cdot \nabla) \nabla \theta = \nabla \times \mathbf{u} \times \nabla \theta - (\nabla \theta \cdot \nabla) \mathbf{u}, \quad (10)$$

214 which can be obtained by replacing ω in Eq. (2) by θ . Like
 215 Eq. (2), the effect of the first term on the right-hand side of
 216 Eq. (10) is to rotate $\nabla \theta$ without changing its magnitude. The
 217 amplification of $\nabla \theta$ is due solely to the stretching term
 218 $(\nabla \theta \cdot \nabla) \mathbf{u}$ and is governed by

$$\partial_t |\nabla \theta| + (\mathbf{u} \cdot \nabla) |\nabla \theta| = - \frac{\nabla \theta}{|\nabla \theta|} \cdot (\nabla \theta \cdot \nabla) \mathbf{u} \leq |\nabla \mathbf{u}| |\nabla \theta|.$$

219

220 Equation (11) implies that following the fluid motion, $|\nabla \theta|$
 221 can grow exponentially in time with an instantaneous rate
 222 bounded from above by $|\nabla \mathbf{u}|$. Hence, the behavior of $|\nabla \mathbf{u}|$
 223 holds the key to understanding the dynamics of $\nabla \theta$. Evi-
 224 dently, following the trajectory of a fluid “particle” starting
 225 from $\mathbf{x} = \mathbf{x}_0$ at $t = 0$, the growth of $|\nabla \theta|$ is formally constrained
 226 by

$$|\nabla \theta| \leq |\nabla \theta_0| \exp \left\{ \int_0^t |\nabla \mathbf{u}| d\tau \right\}, \quad (12)$$

227

228 where $\theta_0 = \theta(\mathbf{x}_0, 0)$ and the integral is along the trajectory in
 229 question. Hence, on average, the rate r defined by

$$r = \frac{1}{t} \int_0^t |\nabla \mathbf{u}| d\tau \quad (13)$$

230

provides an upper bound for the exponential growth rate of
 $|\nabla \theta|$. Note that for $\alpha = 1$ ($\langle |\nabla \mathbf{u}|^2 \rangle = \langle |\nabla \theta|^2 \rangle$), a double expo-
 nential growth of $|\nabla \theta|$ is allowed but not necessarily implied
 by the preceding equations. Nevertheless, it is interesting to
 note that Ohkitani and Yamada [24] observed such a behav-
 ior in their simulations, thereby, suggesting a negative answer
 to the question of finite-time singularities in the surface
 quasigeostrophic equation. This is consistent with the proof
 of nonexistence of blowup by Córdoba [26].

B. Linear versus nonlinear growth of $\nabla \theta$

Now for a sense of the behavior of r , we consider
 $\langle |\nabla \mathbf{u}|^2 \rangle^{1/2}$, which bounds $\langle |\nabla \mathbf{u}| \rangle$ from above by the Cauchy-
 Schwarz inequality $\langle |\nabla \mathbf{u}| \rangle \leq \langle |\nabla \mathbf{u}|^2 \rangle^{1/2}$. For $\alpha \in [2, 4]$,
 $\langle |\nabla \mathbf{u}|^2 \rangle^{1/2}$ can be estimated in terms of the inviscid invariants
 using the following version of the Hölder inequality (see, for
 example, Sec. 5 of Ref. [14]):

$$\begin{aligned} \langle |\nabla \mathbf{u}|^2 \rangle^{1/2} &\leq \langle (-\Delta)^{\alpha/4} \psi^2 \rangle^{1-2/\alpha} \langle (-\Delta)^{\alpha/2} \psi^2 \rangle^{2/\alpha-1/2} \\ &= E^{1-2/\alpha} Z^{2/\alpha-1/2}. \end{aligned} \quad (14)$$

So $\langle |\nabla \mathbf{u}|^2 \rangle^{1/2}$ is controlled by the inviscid invariants E and Z .
 For $\alpha \notin [2, 4]$, inequality (14) reverses direction. Further-
 more, if an initial distribution of θ is to forever spread out in
 wave-number space, $\langle |\nabla \mathbf{u}|^2 \rangle^{1/2}$ increases without bound for
 this case. This implies that there exist different regimes of α ,
 for which $\nabla \mathbf{u}$ evolves quite differently, and the active scalar
 gradient dynamics can be characteristically distinct. We discuss
 all these regimes in what follows.

For $\alpha < 2$, the divergence of $\langle |\nabla \mathbf{u}|^2 \rangle^{1/2}$ entails an acceler-
 ated growth of $\nabla \theta$ from an exponential one. This is the super-
 linear regime discussed in the introductory section. This
 superlinearity reaches the usual quadratic nonlinearity of 3D
 turbulence at $\alpha = 1$, where $\langle |\nabla \mathbf{u}|^2 \rangle = \langle |\nabla \theta|^2 \rangle$. Hence, the sur-
 face quasigeostrophic and 3D Euler equations are analogous
 in this aspect. However, the analogy appears to be superficial
 as the surface quasigeostrophic equation turns out to be far
 more “manageable” than its 3D counterpart: a consequence
 of the material conservation of θ . For example, a number of
 global regularity results have been proved for the surface
 quasigeostrophic equation, by making use of mild dissipation
 mechanisms represented by $(-\Delta)^\eta$ with $\eta \geq 1/2$ [27–30],
 which can be much weaker than the usual viscosity. Whereas
 for the 3D Navier-Stokes system, viscosity appears to be
 inadequate for the same purpose. For $\alpha < 1$, this quadratic
 nonlinearity is surpassed as the ratio $\langle |\nabla \mathbf{u}|^2 \rangle / \langle |\nabla \theta|^2 \rangle$ diverges
 in the limit $\langle |\nabla \theta|^2 \rangle \rightarrow \infty$ because

$$\langle |\nabla \theta|^2 \rangle^{2-\alpha} \leq \langle |\nabla \mathbf{u}|^2 \rangle \langle \theta^2 \rangle^{1-\alpha} \quad (15)$$

(cf. Ref. [14]). Active scalar gradient production can then
 become highly intense.

For $\alpha \in [2, 4]$, $\nabla \mathbf{u}$ is well behaved in the sense that its
 mean square is bounded from above in terms of the inviscid
 invariants [see Eq. (14)]. In this case, $\nabla \mathbf{u}$ is virtually unaf-
 fected by the direct transfer of $\langle \theta^2 \rangle$. At large t , a general fluid

282 trajectory is likely to have traversed the domain many times.
 283 The time average in Eq. (13) may therefore be approximately
 284 replaced by the spatial average. Hence, we can write

$$285 \quad r \approx \langle |\nabla \mathbf{u}| \rangle \leq \langle |\nabla \mathbf{u}|^2 \rangle^{1/2} \leq E^{1-2/\alpha} Z^{2/\alpha-1/2}, \quad (16)$$

286 where we have used the Cauchy-Schwarz inequality and Eq.
 287 (14). This approximation of r means that $\nabla \theta$ can grow expo-
 288 nentially in time without acceleration. Thus, approximately
 289 linear small-scale dynamics can be expected. Note that θ
 290 behaves almost as a passive scalar in this regime. The anal-
 291 ogy between this case and that of a passive scalar was sug-
 292 gested by Schorghofer [12] on phenomenological grounds.

293 When $\alpha > 4$, inequality (14) reverses direction, and
 294 $\langle |\nabla \mathbf{u}|^2 \rangle^{1/2}$ can no longer be controlled by the inviscid invari-
 295 ants. However, unlike the case $\alpha < 2$, for which $\langle |\nabla \mathbf{u}|^2 \rangle^{1/2}$
 296 diverges toward small scales, when $\alpha > 4$ velocity gradients
 297 can be produced at increasingly large scales only. This pro-
 298 duction depends on the inverse transfer of the generalized
 299 energy E (Tran 2004). Within the direct transfer range, i.e.,
 300 the generalized enstrophy range, the portion of $\langle |\nabla \mathbf{u}|^2 \rangle$, say
 301 Ω , cannot increase and instead remains bounded from above
 302 in terms of Z . More precisely, as the spectra of $\langle |\nabla \mathbf{u}|^2 \rangle$ and Z
 303 differ by the factor $k^{2\alpha-4}$, we have $\Omega \leq 2k_*^{4-2\alpha} Z$ (Poincaré
 304 type inequality), where k_* is the lower wave-number end of
 305 the generalized enstrophy range. This suggests that no sig-
 306 nificant changes in the effective degree of nonlinearity of the
 307 small-scale dynamics occur when α exceeds 4. Thus, we can
 308 expect approximately linear small-scale behavior for all α
 309 ≥ 2 .

310 In passing, it is worth mentioning that while the small-
 311 scale dynamics appear to be insensitive to α in the regime
 312 $\alpha > 2$, the large-scale dynamics can vary dramatically. The
 313 reason is that for large α , \mathbf{u} is prone to divergence toward
 314 large scales as the inverse transfer of E proceeds. This un-
 315 doubtedly intensifies motions at large scales. One may adapt
 316 the present notion of degree of nonlinearity for a quantitative
 317 measure of the large-scale dynamics. Analogous to the tradi-
 318 tional problem of regularity, which is concerned with the
 319 possible divergence of $\nabla \theta$, there is a potential problem that \mathbf{u}
 320 becomes divergent for sufficiently large α if the fluid is un-
 321 bounded. This interesting problem is left for a future study.

322 IV. LOCALITY OF THE SMALL-SCALE DYNAMICS

323 This section is concerned with the small-scale dynamics
 324 at the modal level. We establish a connection between the
 325 degree of nonlinearity and dynamical activity of typical local
 326 triads at small scales. Here the dynamical activity of a given
 327 triad is associated with the magnitude of the coupling coef-
 328 ficients within the triad and is independent of the amplitude
 329 of the three modal members. These local triads are shown to
 330 be highly active for $\alpha < 2$ and moderately active for $\alpha = 2$ but
 331 become virtually inactive for $\alpha > 2$. This implies that higher
 332 effective degrees of nonlinearity correspond to more dynami-
 333 cally intense local triads. Thus, the effective degree of non-
 334 linearity is also a measure of dynamical activity of local
 335 triads at small scales. The transition at $\alpha = 2$ from high activ-
 336 ity to virtually no activity of local triads is consistent with
 337 phenomenological arguments [11] that the generalized en-

strophy cascade is spectrally local for $\alpha < 2$ but becomes 338
 dominated by nonlocal interactions for $\alpha > 2$. Below, we also 339
 examine the dynamics of nonlocal triads and elaborate on the 340
 nature of the locality transition, in order to provide a detailed 341
 picture of the direct transfer of $\langle \theta^2 \rangle$ at the modal level. 342

Within each individual triad $\mathbf{k} = \ell + \mathbf{m}$, the transfer of 343
 modal generalized enstrophy is governed by 344

$$\frac{d}{dt} |\hat{\theta}(\mathbf{k})|^2 = \frac{(m^\alpha - \ell^\alpha) \ell \times \mathbf{m}}{m^\alpha \ell^\alpha} [\hat{\theta}(\ell) \hat{\theta}(\mathbf{m}) \hat{\theta}^*(\mathbf{k})$$

$$+ \hat{\theta}^*(\ell) \hat{\theta}^*(\mathbf{m}) \hat{\theta}(\mathbf{k})] = C_k [\hat{\theta}(\ell) \hat{\theta}(\mathbf{m}) \hat{\theta}^*(\mathbf{k})$$

$$+ \hat{\theta}^*(\ell) \hat{\theta}^*(\mathbf{m}) \hat{\theta}(\mathbf{k})], \quad (17) \quad 347$$

$$\frac{d}{dt} |\hat{\theta}(\ell)|^2 = \frac{(k^\alpha - m^\alpha) \ell \times \mathbf{m}}{k^\alpha m^\alpha} [\hat{\theta}(\mathbf{k}) \hat{\theta}^*(\mathbf{m}) \hat{\theta}^*(\ell)$$

$$+ \hat{\theta}^*(\mathbf{k}) \hat{\theta}(\mathbf{m}) \hat{\theta}(\ell)] = C_\ell [\hat{\theta}(\mathbf{k}) \hat{\theta}^*(\mathbf{m}) \hat{\theta}^*(\ell)$$

$$+ \hat{\theta}^*(\mathbf{k}) \hat{\theta}(\mathbf{m}) \hat{\theta}(\ell)], \quad 350$$

$$\frac{d}{dt} |\hat{\theta}(\mathbf{m})|^2 = \frac{(\ell^\alpha - k^\alpha) \ell \times \mathbf{m}}{\ell^\alpha k^\alpha} [\hat{\theta}(\mathbf{k}) \hat{\theta}^*(\ell) \hat{\theta}^*(\mathbf{m})$$

$$+ \hat{\theta}^*(\mathbf{k}) \hat{\theta}(\ell) \hat{\theta}(\mathbf{m})] = C_m [\hat{\theta}(\mathbf{k}) \hat{\theta}^*(\ell) \hat{\theta}^*(\mathbf{m})$$

$$+ \hat{\theta}^*(\mathbf{k}) \hat{\theta}(\ell) \hat{\theta}(\mathbf{m})], \quad (17) \quad 353$$

where we have used the identities $\ell \times \mathbf{m} = \ell \times \mathbf{k} = \mathbf{k} \times \mathbf{m}$ and 354
 suppressed the time variable. It is well known that both E 355
 and Z are conserved for each individual triad. This can be 356
 readily verified by the fact that the coupling coefficients in 357
 Eqs. (17), C_k , C_ℓ and C_m , satisfy 358

$$C_k + C_\ell + C_m = 0 = \frac{C_k}{k^\alpha} + \frac{C_\ell}{\ell^\alpha} + \frac{C_m}{m^\alpha}. \quad 359$$

Furthermore, the transfer of E and Z is from the intermediate 360
 wave number to both the larger and smaller wave numbers or 361
 vice versa (note the signs of the coupling coefficients). The 362
 former behavior appears to have been observed in numerical 363
 simulations of 2D turbulence without exception. 364

We now analyze the coupling coefficients C_k , C_ℓ and C_m 365
 in detail. As crude estimates that hold in general, these can 366
 be bounded by (assuming $k < l < m$) 367

$$|C_k| = \frac{|(m^\alpha - \ell^\alpha) \ell \times \mathbf{m}|}{m^\alpha \ell^\alpha} < k \ell^{1-\alpha}, \quad 368$$

$$|C_\ell| = \frac{|(k^\alpha - m^\alpha) \ell \times \mathbf{m}|}{k^\alpha m^\alpha} < \ell k^{1-\alpha}, \quad 369$$

$$|C_m| = \frac{|(\ell^\alpha - k^\alpha) \ell \times \mathbf{m}|}{\ell^\alpha k^\alpha} < \ell k^{1-\alpha}. \quad (18) \quad 370$$

where we have used $|\ell \times \mathbf{m}| = |\ell \times \mathbf{k}| \leq k \ell$. Similar estimates 371
 were obtained in [20] (for $\alpha = 1, 2$) and in [31] (for $\alpha = 1$). For 372
 $\alpha > 2$, local triads (i.e., $k \leq \ell \leq m$) at small scales are effec- 373
 tively “turned off” because all C_k , C_ℓ and C_m tend to zero in 374

375 the limit $k \rightarrow \infty$. Furthermore, the convergence is as rapid as
 376 $k^{2-\alpha}$. An immediate interpretation of this observation is that
 377 local triads can be relatively ineffective in the direct transfer
 378 of $\langle \theta^2 \rangle$ compared with their nonlocal counterparts (see be-
 379 low). At the critical value $\alpha=2$, C_k , C_ℓ and C_m can remain
 380 order unity for local triads that satisfy $|\ell \times m| \approx k^2$ and $|m^\alpha$
 381 $-\ell^\alpha| \approx |k^\alpha - m^\alpha| \approx |\ell^\alpha - k^\alpha| \approx k^\alpha$. A majority of local triads sat-
 382 isfy both of these conditions. They are neither “ultrathin” nor
 383 nearly isosceles and correspond to relatively sharp estimates
 384 in Eqs. (18), which reduce to $|C_k| \approx |C_\ell| \approx |C_m| \approx 1$. This
 385 means that local triads at small scales in the usual vorticity
 386 equation are moderately active. They can play a significant
 387 role in the direct transfer. Finally, for $\alpha < 2$, the interaction
 388 coefficients of these same triads diverge as $k \rightarrow \infty$. Their di-
 389 vergence can be seen to be as rapid as $k^{2-\alpha}$. This result sug-
 390 gests that for this case, local triads can play an overwhelm-
 391 ingly dominant role in the direct transfer.

392 Next, we turn to nonlocal triads. These are thin triads with
 393 the wave numbers k , ℓ and m , satisfying $k \ll \ell \leq m$. For this
 394 case, C_k , C_ℓ and C_m can be estimated as follows:

$$\begin{aligned}
 |C_k| &= \frac{|(m^\alpha - \ell^\alpha)\ell \times m|}{m^\alpha \ell^\alpha} \approx \frac{\alpha k^2}{\ell^\alpha}, \\
 |C_\ell| &= \frac{|(k^\alpha - m^\alpha)\ell \times m|}{k^\alpha m^\alpha} \approx \ell k^{1-\alpha}, \\
 |C_m| &= \frac{|(\ell^\alpha - k^\alpha)\ell \times m|}{\ell^\alpha k^\alpha} \approx \ell k^{1-\alpha}.
 \end{aligned}
 \tag{19}$$

398 In the limit $\ell \rightarrow \infty$ (while $k < \infty$), C_k vanishes, but both C_ℓ
 399 and C_m ($C_\ell \approx -C_m$) diverge as rapidly as ℓ . This implies a
 400 vigorous exchange of generalized enstrophy between the two
 401 neighboring wave numbers ℓ and m , mediated by a virtually
 402 nonparticipating distant wave number k . This ultralocal
 403 transfer by nonlocal interactions is virtually independent of α
 404 as the divergence of C_ℓ and C_m is insensitive to α . This result
 405 implies that local transfer by nonlocal interactions is an in-
 406 trinsic characteristic of this family of 2D turbulence models.
 407 Note, however, that this transfer can be significant only when
 408 the spectrum of the generalized enstrophy is not steeper than
 409 k^{-1} [32]. In other words, the generalized enstrophy needs to
 410 be physically present at small scales in order to facilitate
 411 such a transfer. This suggests that for $\alpha > 2$ (recall that local
 412 triads are dynamically inactive), the generalized enstrophy
 413 spectra can plausibly scale as k^{-1} because steeper spectra are
 414 unable to support a non-negligible direct transfer. This uni-
 415 versal scaling was suggested by Schorghofer [12] and Wan-
 416 tanabe and Iwayama [15]. Their justification is that θ can be
 417 considered as a passive scalar, a view in accord with the
 418 present analysis.

419 In passing, it is worth mentioning that the divergence of
 420 C_ℓ and C_m in nonlocal triads is probably the reason for nu-
 421 merical instability in simulations of 2D turbulence with in-

adequate diffusion because local triads with coupling coeffi- 422
 cients of order unity are evidently well behaved. Support for 423
 this claim can be derived from common observations that 424
 numerical divergences occur as soon as the modes in the 425
 vicinity of the truncation wave number are excited and well 426
 before they acquire any considerable amount of enstrophy. 427
 The same instability problem persists for $\alpha > 2$, although the 428
 weak activities of local triads in this case may reduce the 429
 severity of the instability to a certain extent. 430

V. CONCLUDING REMARKS 431

We have presented the notion of effective degree of non- 432 AQ:
 linearity to quantify the small-scale dynamics of a family of 433 #1
 generalized models of two-dimensional turbulence governed 434
 by a broad class of nonlinear transport equations. Here, the 435
 active scalar $\theta = (-\Delta)^{\alpha/2} \psi$ ($\alpha > 0$) is advected by the incom- 436
 pressible flow $\mathbf{u} = (-\psi_y, \psi_x)$, where ψ is the stream function. 437
 We have argued that although the advection term is qua- 438
 dratic, the effective degree of nonlinearity of the small-scale 439
 dynamics is not always quadratic and depends on α . It has 440
 been found that the active scalar gradient dynamics are vir- 441
 tually linear for $\alpha \geq 2$ and become nonlinear for $\alpha < 2$. Fur- 442
 thermore, the degree of nonlinearity increases as α is de- 443
 creased from 2, becoming quadratic at $\alpha=1$ and exceeding 444
 quadratic nonlinearity for $\alpha < 1$. It is conceivable that cred- 445
 ible theories of the family’s dynamics, particularly, those in- 446
 volving small scales, need to account for the dependence on 447
 α of the effective degree of nonlinearity. 448

We have also found that local triads at small scales are 449
 highly active for $\alpha < 2$, moderately active for $\alpha=2$, and vir- 450
 tually inactive for $\alpha > 2$. On the other hand, nonlocal triads 451
 are characterized by a vigorous exchange of generalized en- 452
 strophy between pairs of neighboring wave numbers, medi- 453
 ated by the third nonparticipating distant wave number. This 454
 property is common for all α , thereby, implying that nonlocal 455
 interactions (but ultralocal transfer) can be considered uni- 456
 versal. In the absence of local triad activity ($\alpha > 2$), this ul- 457
 tralocal transfer is responsible for the direct transfer of gene- 458
 ralized enstrophy. This is similar to the problem of passive 459
 scalar transport by a large-scale flow as the weak feedback 460
 on the advecting flow by the active scalar can be neglected 461
 [32]. In this case, it appears plausible that generalized enstro- 462
 phy spectra scale as k^{-1} . 463

The local nature of the generalized enstrophy transfer can 464
 be seen to be unambiguous in the present study. In general, 465
 this transfer is local in wave-number space regardless of 466
 what types of triads make the most contribution. For local 467
 triads, the generalized enstrophy transfer is inherently local. 468
 For nonlocal interactions, the transfer is even “more” local, 469
 having a relatively higher degree of locality compared to the 470
 transfer by local triads. More importantly, the transfer be- 471
 tween distant wave numbers is largely insignificant. Hence, it 472
 makes sense to speak of the degree of locality of the direct 473
 generalized enstrophy transfer rather than to distinguish be- 474
 tween local and distant transfers. 475

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- 478** [1] R. K. Kraichnan, Phys. Fluids **10**, 1417 (1967).
479 [2] R. H. Kraichnan, J. Fluid Mech. **47**, 525 (1971).
AQ: 480 [3] G. K. Batchelor, Phys. Fluids **12**, II-233 (1969).
#2 481 [4] J. Charney, J. Atmos. Sci. **28**, 1087 (1971).
482 [5] R. M. Kerr, Phys. Fluids A **5**, 1725 (1993).
483 [6] R. M. Kerr, Phys. Fluids **17**, 075103 (2005).
484 [7] P. Dmitruk and D. C. Montgomery, Phys. Fluids **17**, 035114
485 (2005).
486 [8] C. V. Tran and D. G. Dritschel, J. Fluid Mech. **559**, 107
487 (2006).
488 [9] D. G. Dritschel, C. V. Tran, and R. K. Scott, J. Fluid Mech.
489 591, 379 (2007).
AQ: 490 [10] C. V. Tran, Phys. Fluids (to be published).
#3 491 [11] R. T. Pierrehumbert, I. M. Held, and K. L. Swanson, Chaos,
492 Solitons Fractals 4, 1111 (1994).
493 [12] N. Schorghofer, Phys. Rev. E **61**, 6572 (2000).
494 [13] K. S. Smith, G. Boccaletti, C. C. Henning, I. Marinov, C. Y.
495 Tam, I. M. Held, and G. K. Vallis, J. Fluid Mech. **469**, 13
496 (2002).
497 [14] C. V. Tran, Physica D **191**, 137 (2004).
498 [15] T. Watanabe and T. Iwayama, J. Phys. Soc. Jpn. **73**, 3319
AQ: 499 (2004).
#4 500 [16] I. M. Held, R. T. Pierrehumbert, S. T. Garner, and K. L. Swan-
son, J. Fluid Mech. **282**, 1 (1995). **501**
[17] R. K. Scott, Phys. Fluids **18**, 116601 (2006). **502**
[18] C. V. Tran, Physica D **213**, 76 (2006). **503**
[19] C. V. Tran and J. C. Bowman, J. Fluid Mech. **526**, 349 (2005). **504**
[20] C. V. Tran and D. G. Dritschel, Phys. Fluids **18**, 121703 **505**
(2006). **506**
[21] P. Constantin, A. Majda, and E. Tabak, Nonlinearity **7**, 1495 **507**
(1994a). **508**
[22] P. Constantin, A. Majda, and E. Tabak, Phys. Fluids **6**, 9 **509**
(1994b). **510**
[23] A. Córdoba and D. Córdoba, Commun. Math. Phys. **249**, 511 **511**
(2004). **512**
[24] K. Ohkitani and M. Yamada, Phys. Fluids **9**, 876 (1997). **513**
[25] A. F. Bennett, AIP Conf. Proc. **106**, 295 (1984). **514**
[26] D. Córdoba, Ann. Math. **148**, 1135 (1998). **515**
[27] J. Carrillo and L. Ferreira, Nonlinearity **21**, 1001 (2008). **516**
[28] P. Constantin, Nonlinearity **21**, T239 (2008). **517 AQ:**
[29] N. Ju, Commun. Math. Phys. **255**, 161 (2005). **518 #5**
[30] A. Kiselev, F. Nazarov, and A. Volberg, Invent. Math. **167**, **519**
445 (2007). **520**
[31] P. Constantin, Physica D **237**, 1926 (2008). **521**
[32] C. V. Tran, Phys. Rev. E **78**, 036310 (2008). **522**

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